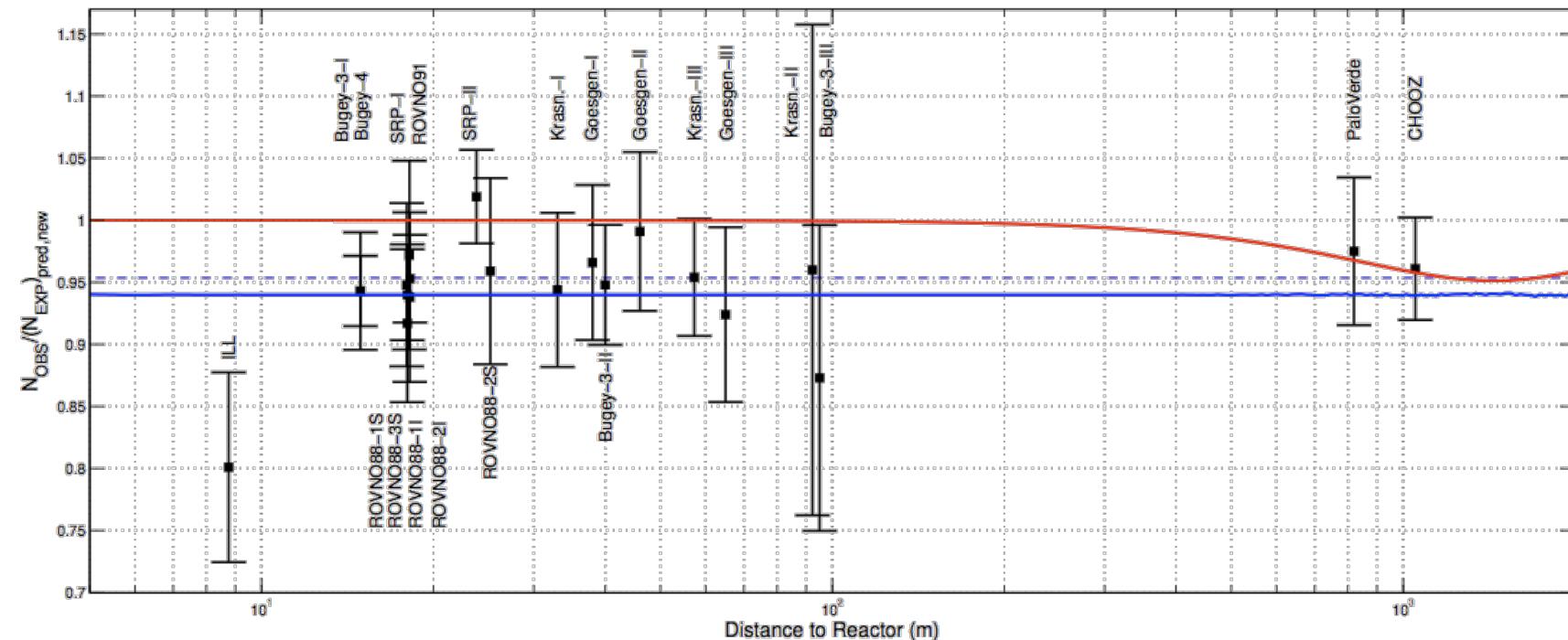


# The Reactor Antineutrino Anomaly and its Implications



G. Mention (CEA-Saclay, Irfu SPP)

# NEW REACTOR ANTINEUTRINO SPECTRA

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*CEA / Irfu & IN2P3 / Subatech*

[arXiv:1101.2663 \[hep-ex\]](https://arxiv.org/abs/1101.2663) , accepted in PRC

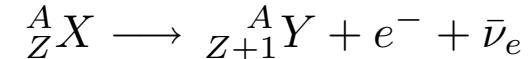
\* corresponding author

# Reactor anti-neutrinos: introduction



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- Electron antineutrinos emitted through decays of **Fission Products** of  $^{235}\text{U}$ ,  $^{238}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$



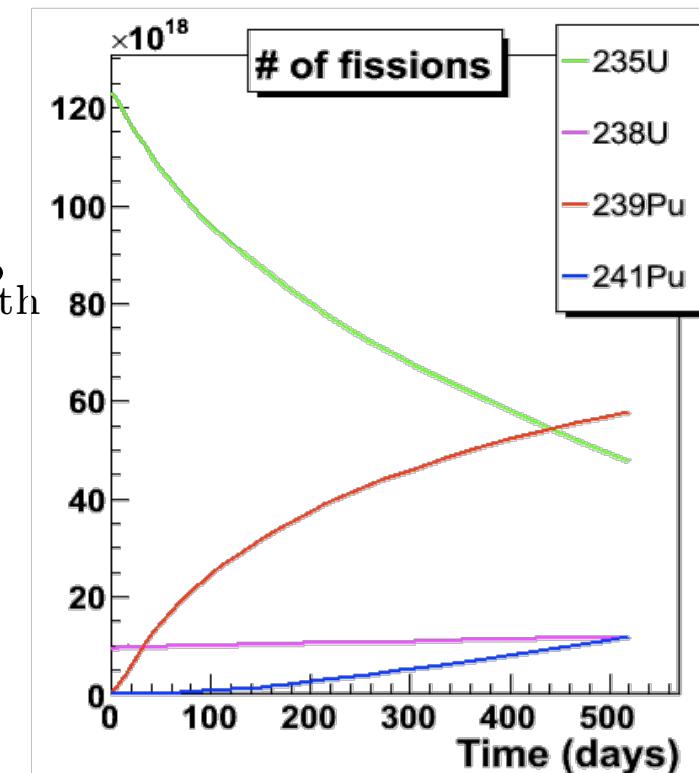
- Each fission product may have **many decay modes or branches**
- About **6 anti- $\nu_e$**  emitted per **fission**.
- Nuclear reactors  $1 \text{ GW}_{\text{th}} \Leftrightarrow 2 \times 10^{20} \bar{\nu}/\text{s}$
- Neutrino Luminosity  $N_{\bar{\nu}} = \gamma(1 + k)P_{\text{th}}$

$N_{\bar{\nu}}$ : neutrino flux

$P_{\text{th}}$ : thermal power

$\gamma$ : reactor constant

$k$ : fuel evolution correction (up to 10%)



# $\nu$ spectrum emitted by a reactor



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The prediction of reactor  $\nu$  spectrum is the dominant source of systematic error for single detector reactor neutrino experiments

## Reactor data

Thermal power,  $\delta P_{th} \leq 1\%$

$$\Phi_\nu(E,t) = \frac{P_{th}(t)}{\sum_k \alpha_k(t) E_k} \times \sum_k \alpha_k(t) S_k(E)$$

## Nuclear databases

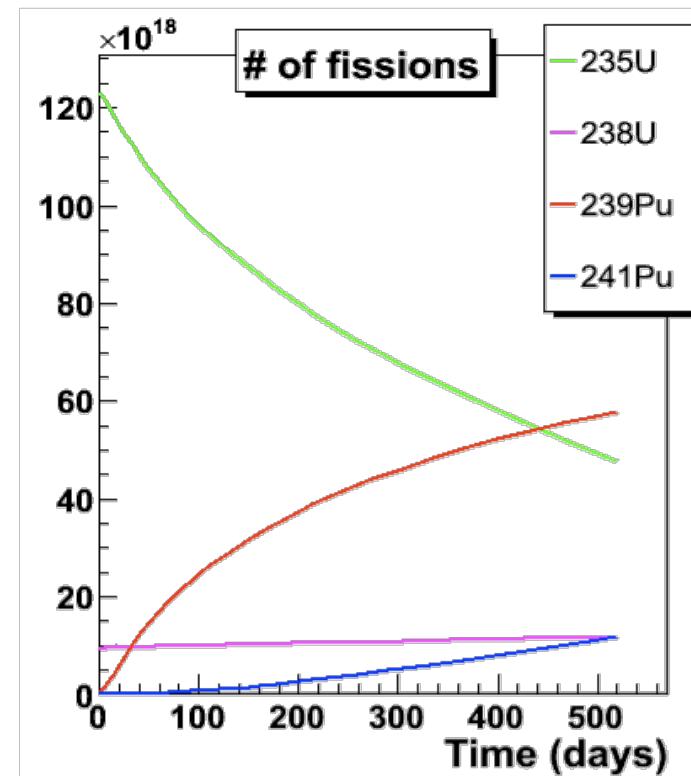
$E$  released per fissions of isotope  $k$ ,  
 $\delta E_k \approx 0.3\%$

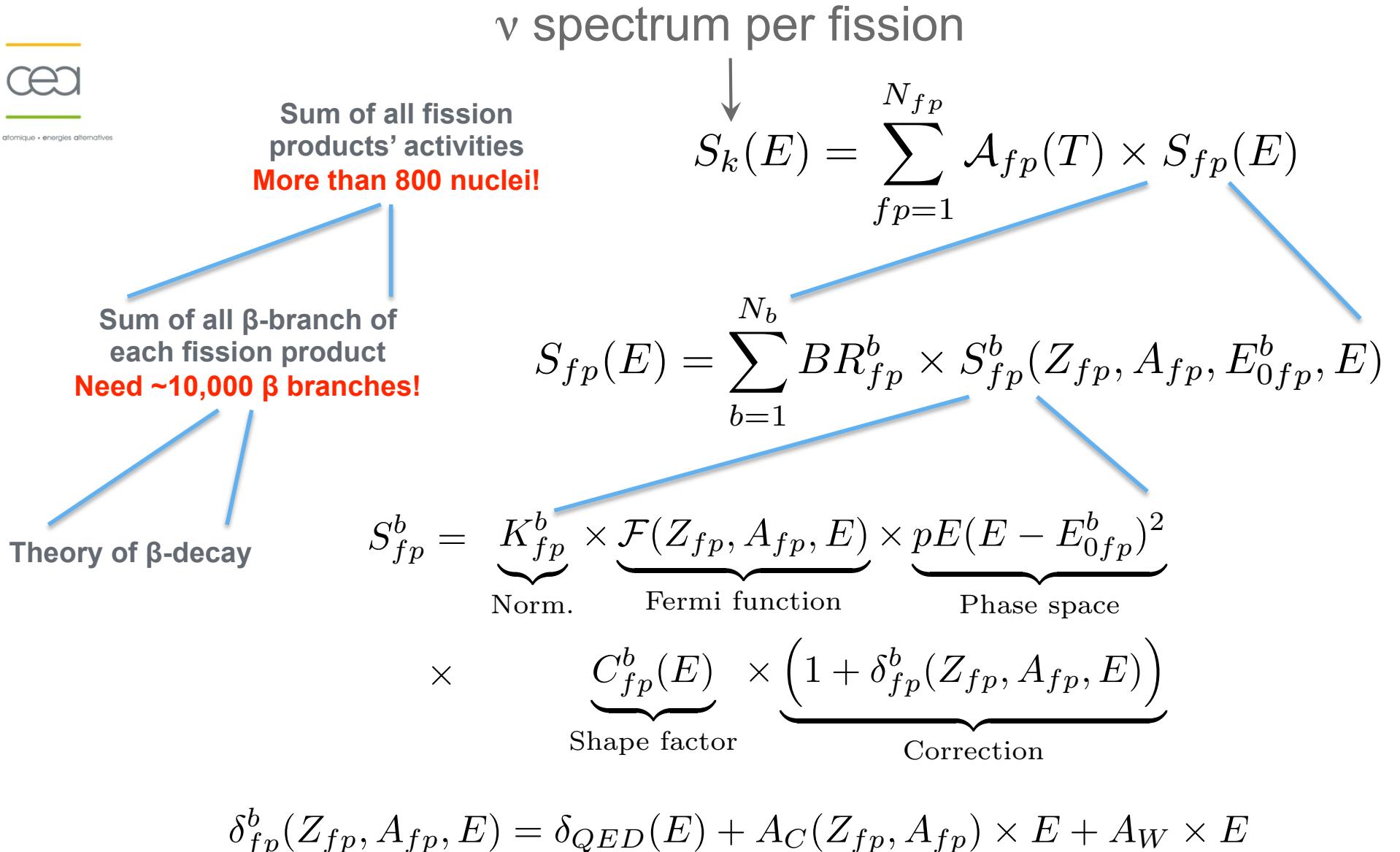
## Reactor evolution codes

Fraction of fissions from isotope  $k$ ,  $\delta a_k = \text{few \%}$   
but large anti-correl @ fixed  $P_{th}$

$\nu$  spectrum per fission  
**This work !**

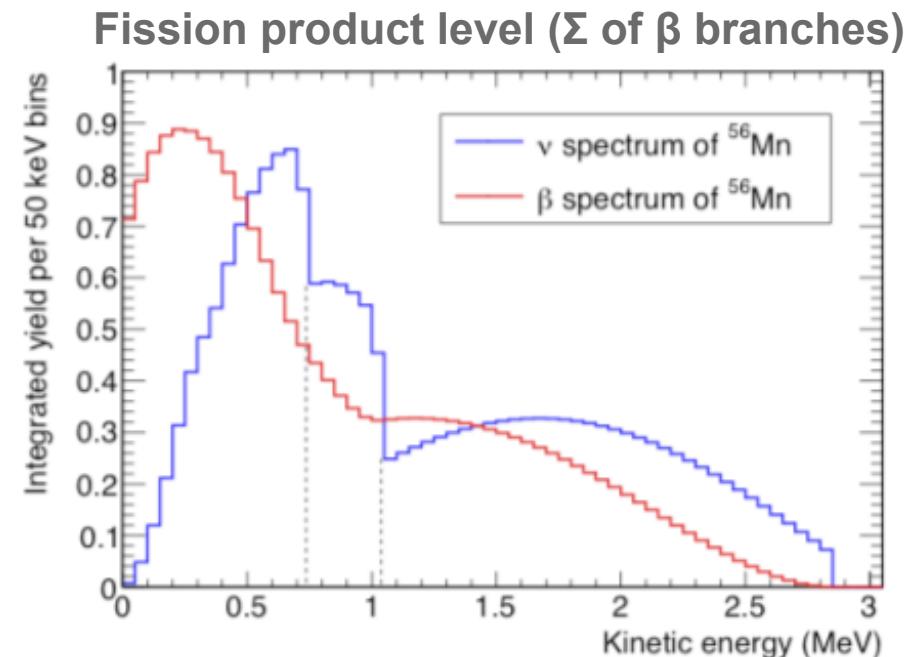
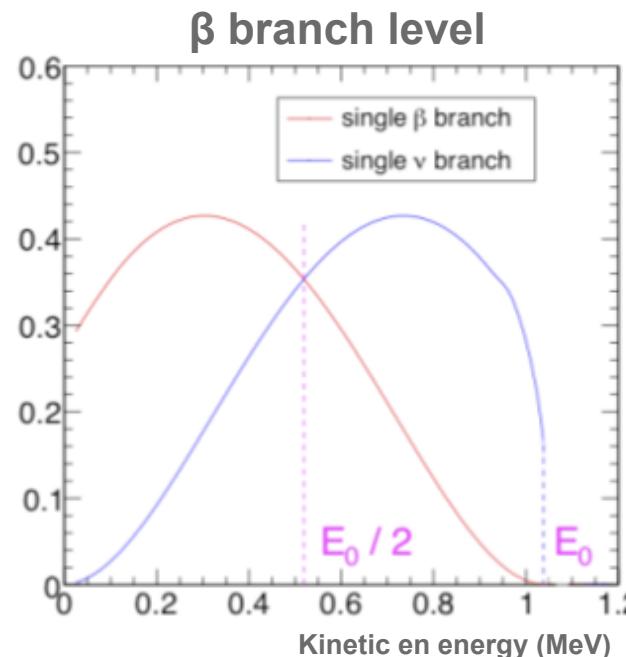
$k = {}^{235}\text{U}, {}^{238}\text{U}, {}^{239}\text{Pu}, {}^{241}\text{Pu}$





# From $e^-$ to anti- $\nu_e$ spectra

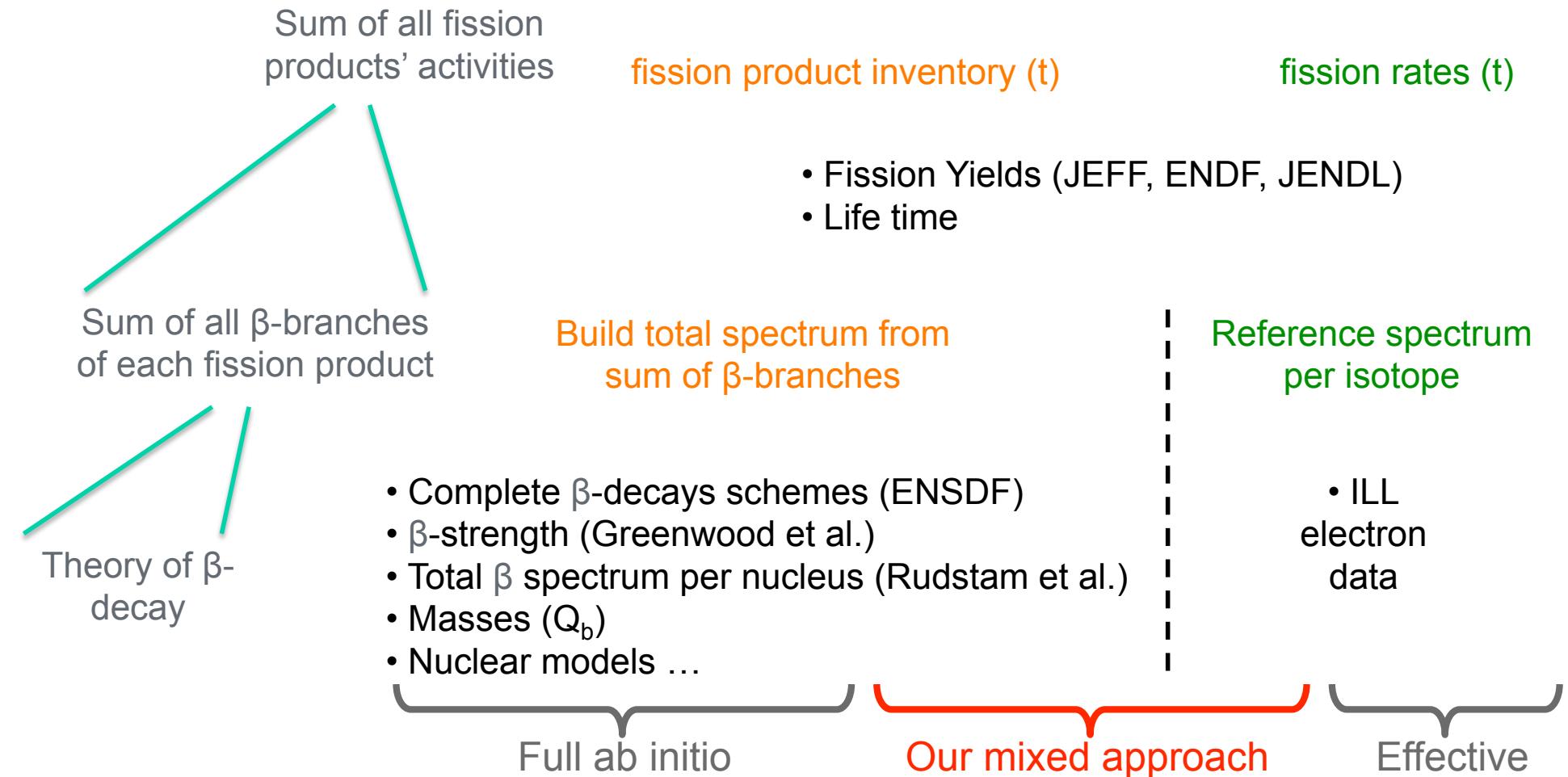
- **A single beta decay branch:**  ${}_Z^A X \longrightarrow {}_{Z+1}^A Y + e^- + \bar{\nu}_e$ 
  - depends on: nucleus (Z), branching ratio (BR), end point (Q), spin-parity
  - Energy conservation:  $E_e + E_\nu = Q_e$
- Anti- $\nu$  spectra are computed from electron spectra by “inverting” each branch separately
- Cannot go from  $e^-$  to  $\nu$  from a global  $e^-$  spectrum, need each individual branch from each contributing nucleus



# Complementary approaches to compute the $\nu$ flux



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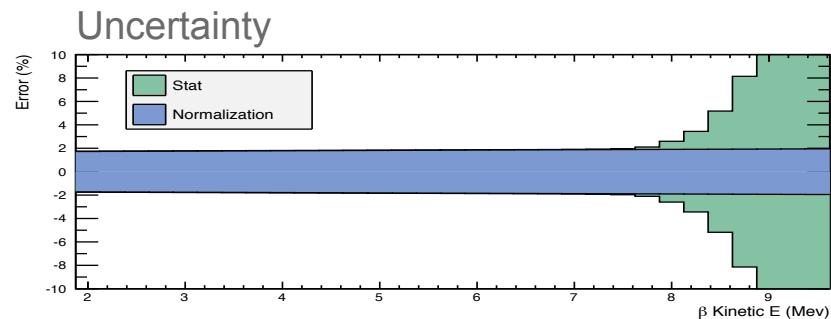
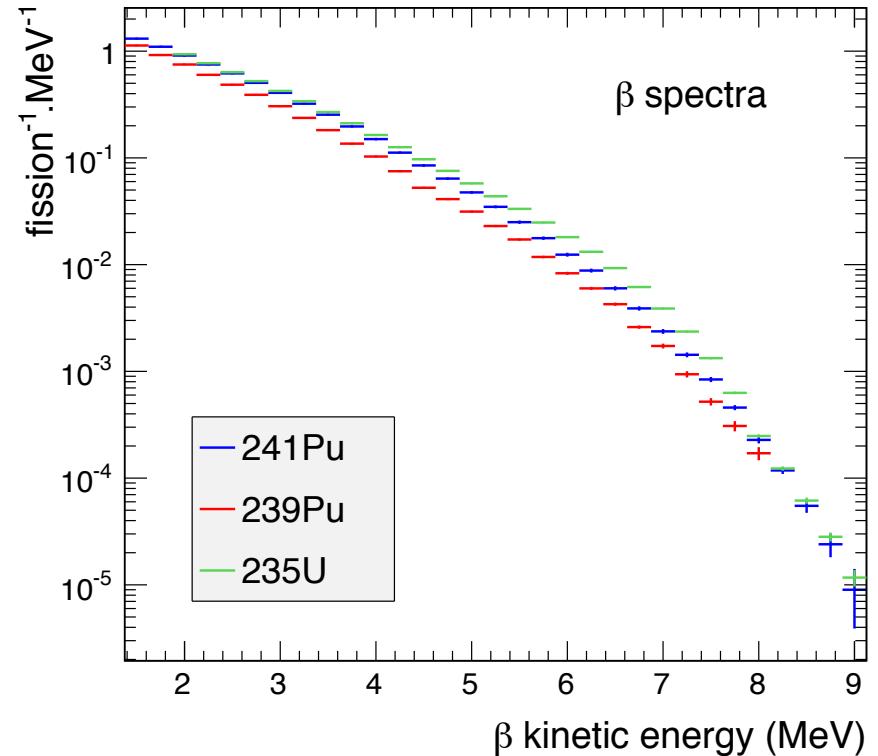
# $\nu$ flux prediction: Anchor point of ILL electron data

Unique reference to be met by any other measurement or calculation

- Accurate measurements @ ILL in Grenoble (1980-89):
  - High resolution magn. spectrometer
  - Intense and pure thermal n spectrum from the core (not suitable for  $^{238}\text{U}$  which needs fast n)
- Measure total  $e^-$  spectrum from decays of fission product.

$$\sum_{A,Z} \left\{ {}_Z^AX \longrightarrow {}_{Z+1}^AY + e^- + \bar{\nu}_e \right\}$$

- Calibration through extensive use of reference internal conversion electron lines
- Normalization syst. @ 1.8%

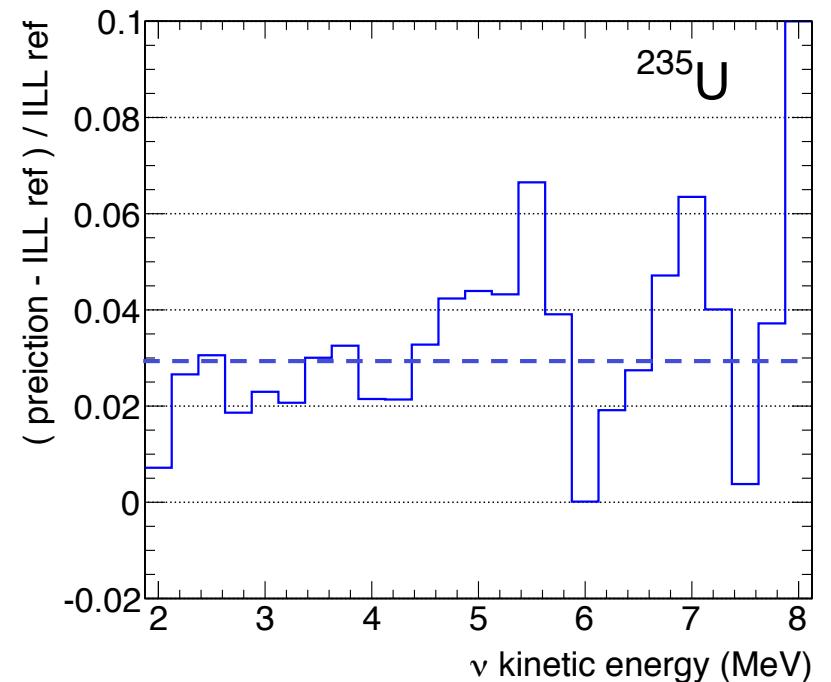
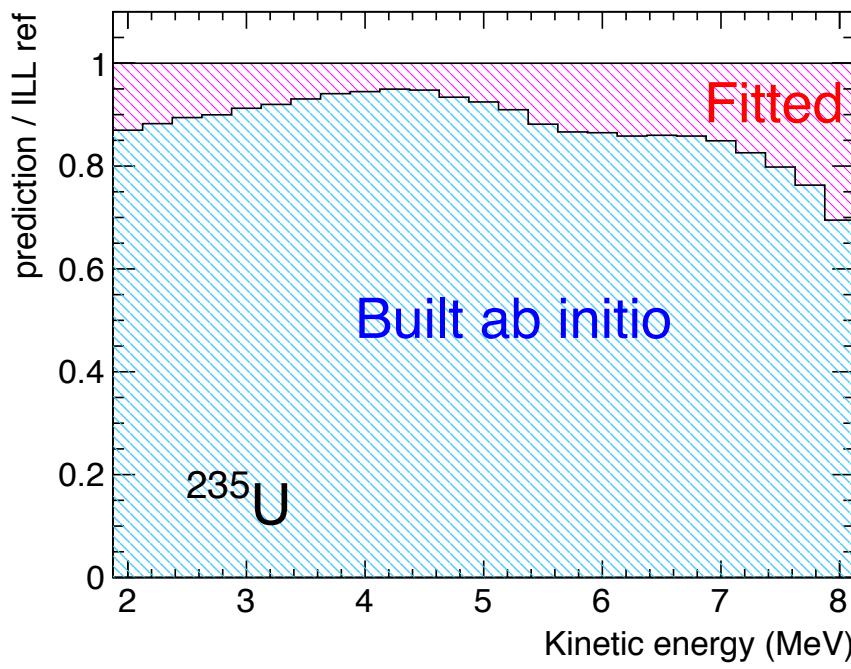


# The New Mixed Conversion Approach



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1. SAME ILL  $e^-$  data Anchorage
2. Ab-Initio: “true” distribution of  $\beta$ -branches reproduces >90% of ILL  $e^-$  data.
3. Old-procedure: reduce use of effective anchorage-branches to the remaining 10%.



- About +3% normalization shift with respect to old  $v$  spectrum
- Similar result for all isotopes ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ )
- Stringent Test Performed – Origin of the bias identified

# Off-equilibrium corrections to ILL $\beta$ -spectra measurements



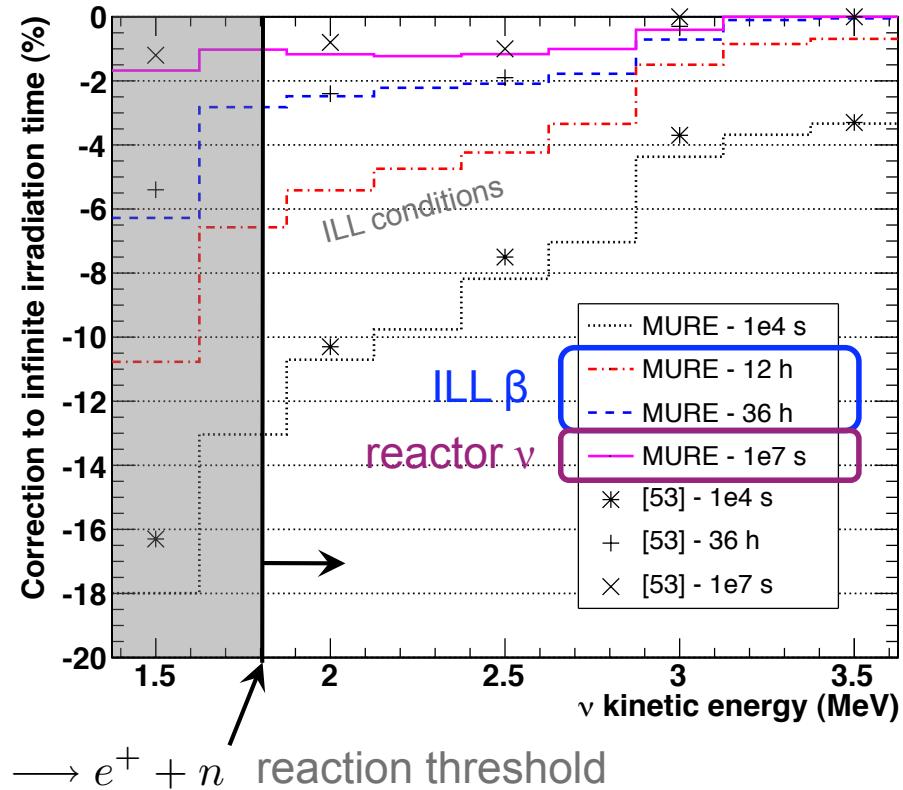
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MURE evolution code: core composition and off equilibrium effects  
(Subatech Nantes)

$$S_k(E) = \sum_{fp=1}^{N_{fp}} \mathcal{A}_{fp}(T) \times S_{fp}(E)$$

- Full simulation of reactor core  
→ absolute prediction of isotopes inventory.
- Relative off-equilibrium effect: close to beta-inverse threshold, **a significant fraction of the  $\nu$  spectrum takes weeks to reach equilibrium**  
→ Sizeable correction in the  $\nu$  oscillation range that depends on the exact chronology of ILL data taking.

Relative change of  $\nu$  spectrum  
w.r.t. infinite irradiation time



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# THE REACTOR ANTINEUTRINO ANOMALY

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M. Cribier, Th. Mueller D. Lhuillier, A. Letourneau,*

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[arXiv:1101.2755 \[hep-ex\]](https://arxiv.org/abs/1101.2755), [PRD83, 073006 \(2011\)](https://doi.org/10.1103/PRD.83.073006)

\* corresponding author

- **Inverse Beta Decay:**  $\bar{\nu}_e + p \rightarrow e^+ + n$
- **Theoretical predictions: our results agree with**
  - Vogel 1984 (Phys Rev D29 p1918)
  - Fayans 1985 (Sov J Nucl Phys 42)
  - Vogel-Beacom 1999: “supersedes” Vogel 84 (Phys Prev D60 053003)

$$\sigma_{V-A}(E_e) = \kappa p_e E_e (1 + \delta_{rec} + \delta_{wm} + \delta_{rad})$$

- **The pre-factor  $\kappa$  (two pseudo-independent approaches)**

$$\kappa = \frac{G_F^2 \cos^2(\theta_C)}{\pi} (1 + \Delta_{inner}^R) (1 + 3\lambda^2) = \frac{2\pi^2}{m_e^5 f^R \tau_n}$$

- **$\kappa$ 's value raised over the history**, from  $0.914 \cdot 10^{-42} \text{ cm}^2$  in 1981
  - Vogel/Beacom 1999 :  $\kappa = 0.952 \cdot 10^{-42} \text{ cm}^2$
  - **Our work is based on 2010 PDG  $\tau_n$**  :  $\kappa = 0.956 \cdot 10^{-42} \text{ cm}^2$
  - But we anticipate 2011  $\kappa = 0.961 \cdot 10^{-42} \text{ cm}^2$  ( $\langle \tau_n \rangle$  revision)

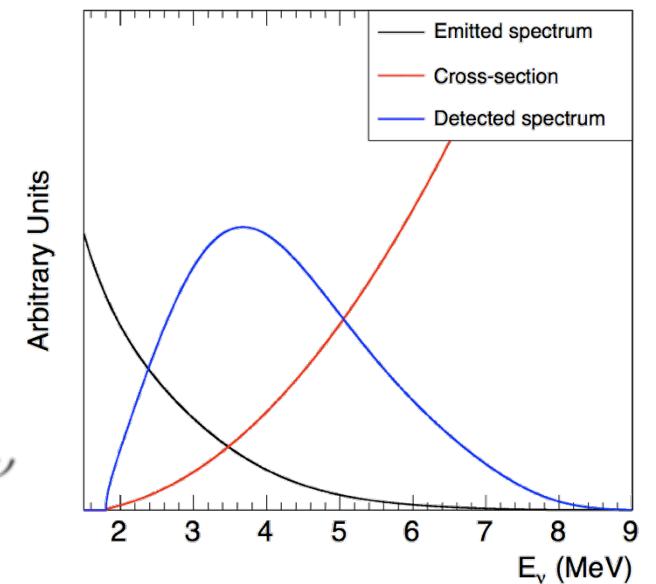
# Reactor electron anti-neutrino detection

- **Inverse Beta Decay:**  $\bar{\nu}_e + p \rightarrow e^+ + n$ 
  - Threshold: 1.806 MeV
- **Anti- $\nu_e$  interaction rate**  $n_\nu = \frac{1}{4\pi R^2} \frac{P_{\text{th}}}{\langle E_f \rangle} N_p \varepsilon \sigma_f$ 
  - $E_{\text{released / fission}}$
  - **Target free protons**
  - **efficiency**
- **Experimental cross section per fission:**  $\sigma_f$

$$\sigma_f^{\text{meas.}} = \frac{4\pi R^2 n_\nu^{\text{meas.}}}{N_p \varepsilon} \frac{\langle E_f \rangle}{P_{\text{th}}}$$

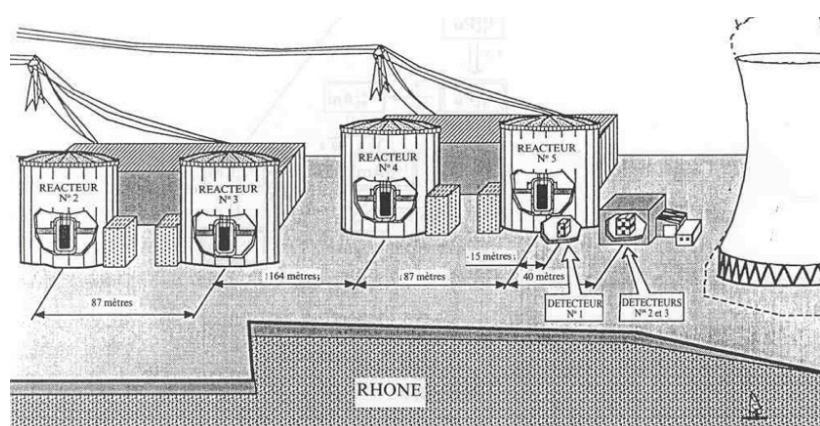
- **Predicted cross section per fission:**  $\sigma_f^{\text{pred.}}$

$$\sigma_f^{\text{pred.}} = \int_0^\infty \phi_f^{\text{pred.}}(E_\nu) \sigma_{\text{V-A}}(E_\nu) dE_\nu$$



# Computing the expected rate/spectrum

$$\sigma_f^{pred} = \int_0^\infty S_{tot}(E_\nu) \sigma_{V-A}(E_\nu) dE_\nu = \sum_k f_k \sigma_{f,k}^{pred}$$



- **Bugey-4 Benchmark**

- Phys. Lett. B 338(1994) 383
- $\tau_n = 887.4$  s
- “old” spectra (30 effective branches)
- no off-equilibrium corrections

$10^{-43} \text{cm}^2/\text{fission}$	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{241}\text{Pu}$
BUGEY-4	$6.39 \pm 1.9\%$	$4.18 \pm 2.4\%$	$5.76 \pm 2.1\%$
This work	$6.39 \pm 1.8\%$	$4.19 \pm 2.3\%$	$5.73 \pm 1.9\%$

- Final agreement to better than 0.1% on best known  $^{235}\text{U}$  w.r.t. their computations
- This validates our calculation code.

# The New Cross Section Per Fission

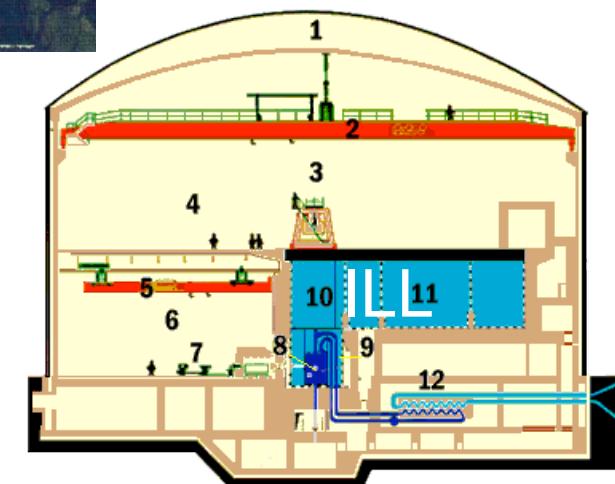
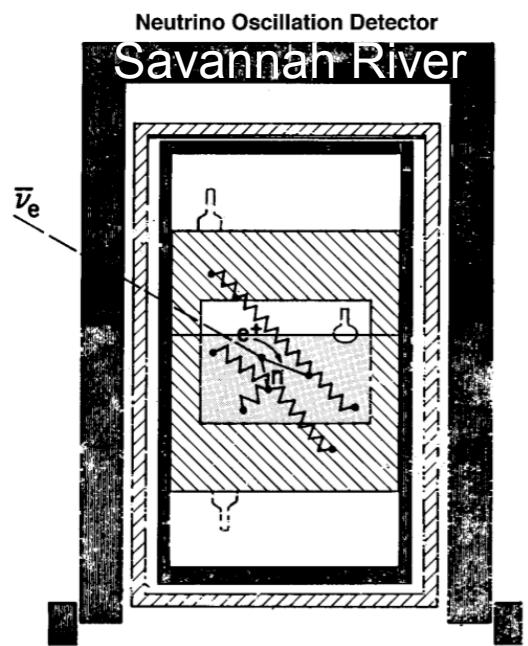
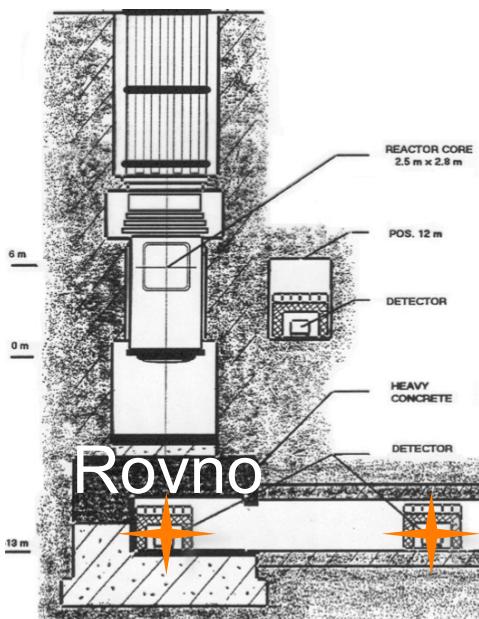
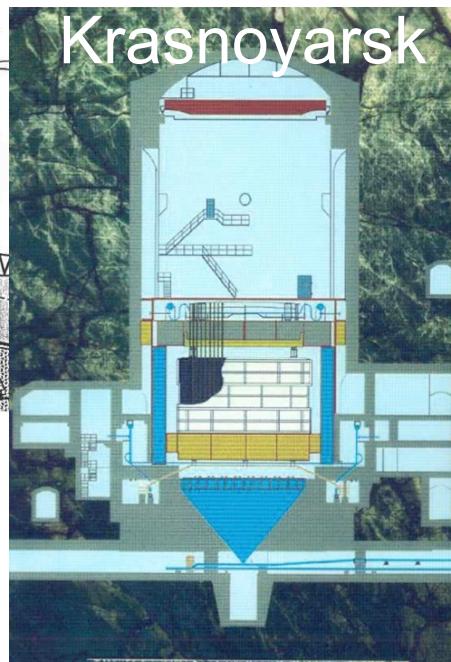
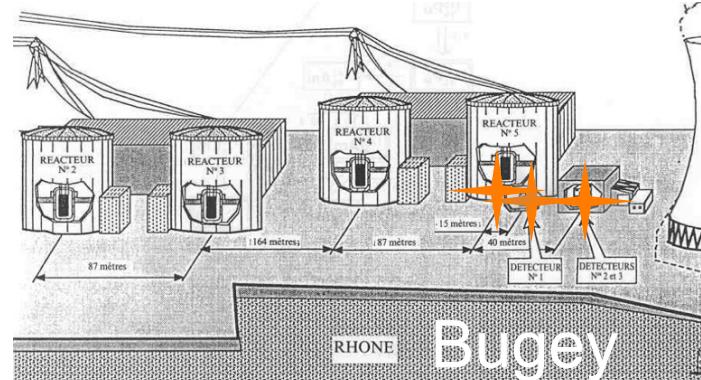
- **v-flux:**  $^{235}\text{U}$  : +2.5%,  $^{239}\text{Pu}$  +3.1%,  $^{241}\text{Pu}$  +3.7%,  $^{238}\text{U}$  +9.8% ( $\sigma_f^{\text{pred}}$  ↗)
- **Off-equilibrium corrections** now included ( $\sigma_f^{\text{pred}}$  ↗)
- **Neutron lifetime** decrease by a few % ( $\sigma_f^{\text{pred}}$  ↗) ( $\sigma_{V-A}(E_\nu) \propto 1/\tau_n$ )
- Slight evolution of the phase space factor ( $\sigma_f^{\text{pred}}$  →)
- Slight evolution of the energy per fission per isotope ( $\sigma_f^{\text{pred}}$  →)
- Burnup dependence:  $\sigma_f^{\text{pred}} = \sum_k f_k \sigma_{f,k}^{\text{pred}}$  ( $\sigma_f^{\text{pred}}$  →)

10 <sup>-43</sup> cm <sup>2</sup> /fission	old [3]	new	relative effect ↓
$\sigma_{f,^{235}\text{U}}^{\text{pred}}$	$6.39 \pm 1.9\%$	$6.61 \pm 2.11\%$	+3.4%
$\sigma_{f,^{239}\text{Pu}}^{\text{pred}}$	$4.19 \pm 2.4\%$	$4.34 \pm 2.45\%$	+3.6%
$\sigma_{f,^{238}\text{U}}^{\text{pred}}$	$9.21 \pm 10\%$	$10.10 \pm 8.15\%$	+9.6%
$\sigma_{f,^{241}\text{Pu}}^{\text{pred}}$	$5.73 \pm 2.1\%$	$5.97 \pm 2.15\%$	+4.2%

# 19 Experimental results at distances below 100 m



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Measured neutrino rates and cross sections per fission  $\sigma_f$

# 19 Experimental Results Revisited ( $L < 100$ m)

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He} + \text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He} + \text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He} + \text{LS}$	889	$\simeq 1$	<0.01	<0.01	<0.01	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He} + \text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He} + \text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

# 19 Experimental Results Revisited (L<100m)

## Neutron lifetime

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
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# 19 Experimental Results Revisited (L<100m)

## Averaged Fuel Composition

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13	SRP I	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He} + \text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He} + \text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

# 19 Experimental Results Revisited ( $L < 100m$ )

OBSERVED/PREDICTED ratios: OLD & NEW (this work)

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He} + \text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	$\text{Li}-\text{LS}$	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He} + \text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He} + \text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He} + \text{LS}$	889	$\simeq 1$	<0.01	<0.01	<0.01	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He} + \text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
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# 19 Experimental Results Revisited (L<100m)

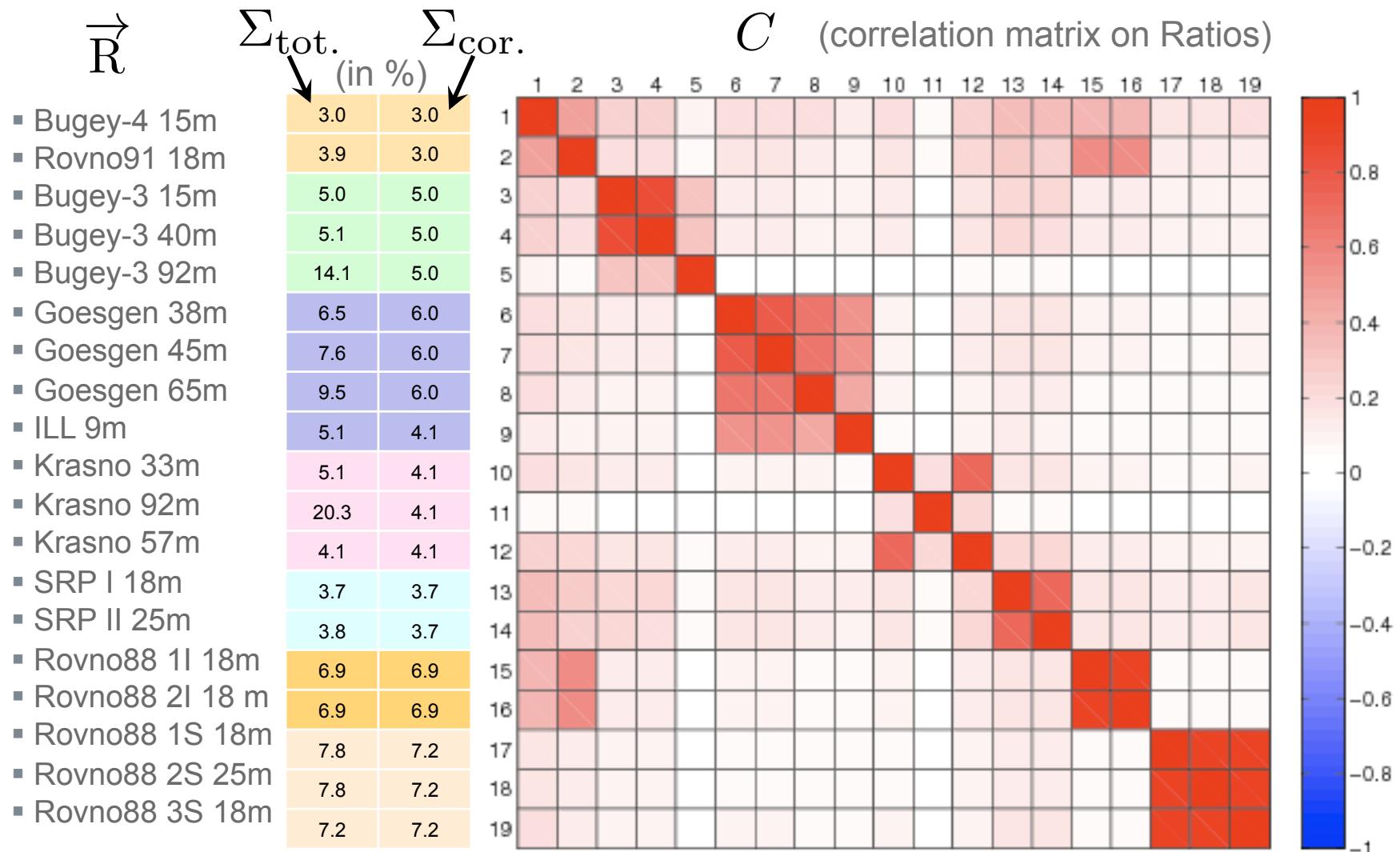
Total Errors Exp.+ $\nu$ -Spectra (%) & Correlated errors (%)



#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
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7	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
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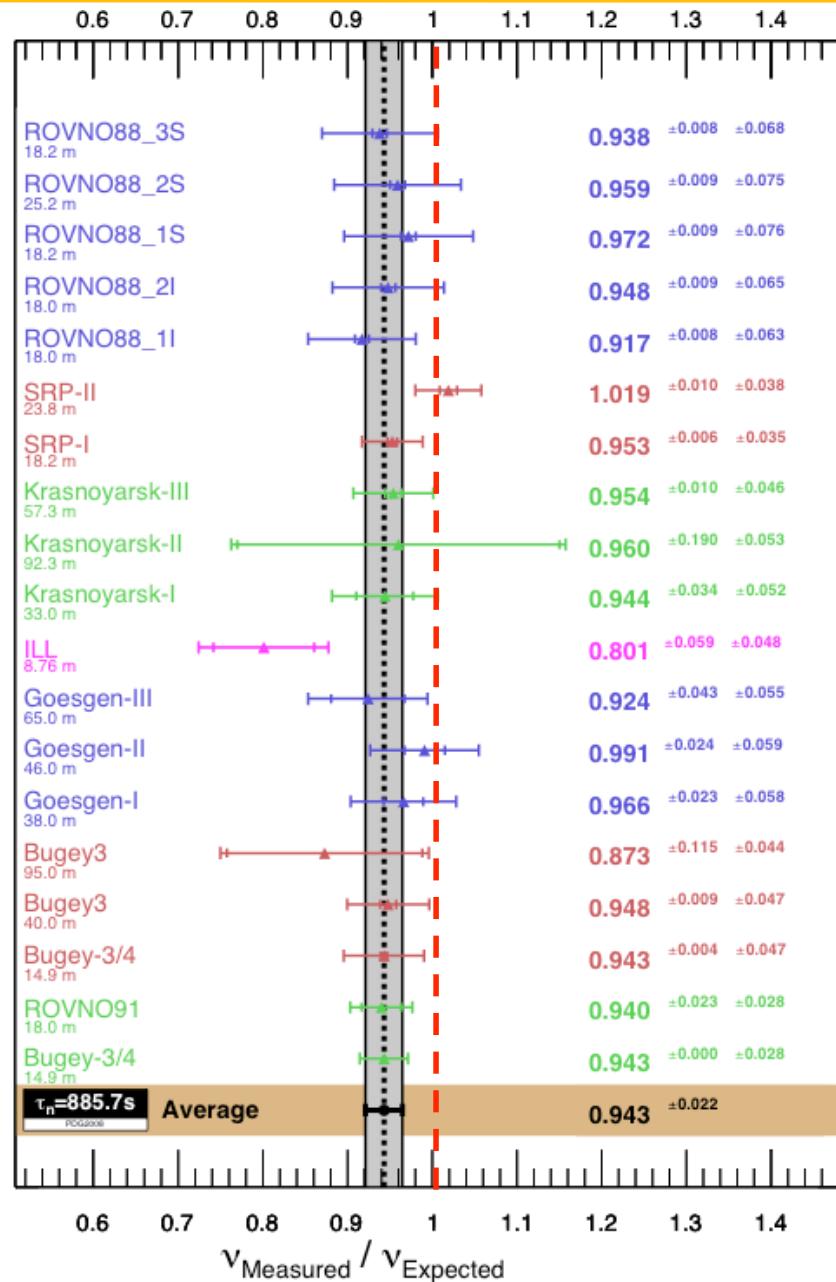
- Our guiding principles: Be conservative & stable numerically
- We correlated experiments in the following way:
  - 2% systematic on flux fully correlated over all measurements of  $\beta$ -spectra of ILL
- Non-flux systematic error correlations across measurements:
  - Same experiment with same technology: 100% correlated
  - ILL shares 6% correlated error with Gösgen although detector slightly different. Rest of ILL error is uncorrelated.
  - Rovno 88 integral measurements 100% corr. with Rovno 91 despite detector upgrade, but not with Rovno 88 LS data
  - Rovno 88 integral meas. 50% correlated with Bugey-4

# Experiments correlation matrix on ratios = meas./pred.



- Main pink color comes from the 2% systematic on ILL  $\beta$ -spectra normalization uncertainty
- The experiment block correlations come from identical detector, technology or neutrino source

# The reactor anti-neutrino anomaly



$$\chi^2 = \left( \mathbf{r} - \overrightarrow{\mathbf{R}} \right)^T W^{-1} \left( \mathbf{r} - \overrightarrow{\mathbf{R}} \right)$$

Weights:  $W = \Sigma_{\text{unc.}}^2 + \Sigma_{\text{cor.}} C \Sigma_{\text{cor.}}$   
with  $\Sigma_{\text{unc.}}^2 = \Sigma_{\text{tot.}}^2 - \Sigma_{\text{cor.}}^2$

The synthesis of published experiments at reactor-detector distances  $\leq 100$  m leads to a ratio R of observed event rate to predicted rate of

$$\mu = 0.976 \pm 0.024 \text{ (OLD flux)}$$

With our **NEW flux** evaluation, this ratio shifts to

$$\mu = 0.943 \pm 0.023,$$

leading to a deviation from unity at 98.6% C.L.

$$\chi^2_{\min} = 19.6/18$$

# The reactor rate anomaly

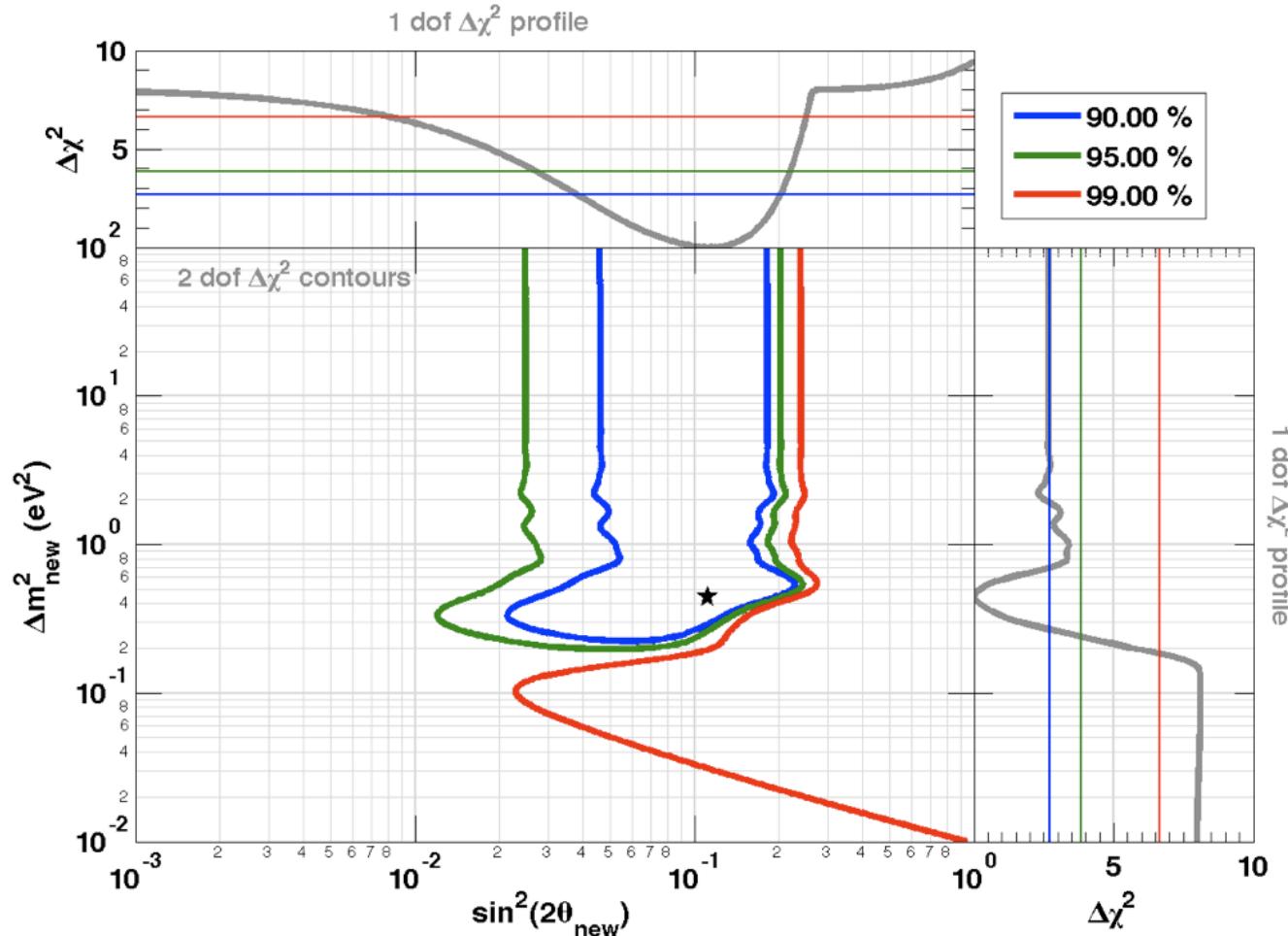
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- 18/19 short baseline experiments <100m from a reactor observed a deficit of anti- $\nu_e$  compared to the new prediction
- The effect is statistically significant at more 98.6%
- Effect partly due to re-evaluation of cross-section parameters, especially updated neutron lifetime, accounting for off equ. effect
- **At least three alternatives:**
  - Our conversion calculations are wrong. Anchorage at the ILL electron data is unchanged w.r.t. old prediction
  - Bias in all short-baseline experiments near reactors : unlikely...
  - New physics at short baselines, explaining a deficit of anti- $\nu_e$  :
    - **Oscillation towards a 4<sup>th</sup>, sterile  $\nu$  ?**
    - **a 4<sup>th</sup> oscillation mode with  $\theta_{\text{new}}$  and  $\Delta m^2_{\text{new}}$**

# The 4<sup>th</sup> neutrino hypothesis

- Combine all rate measurements, no spectral-shape information
- Fit to anti- $\nu_e$  disappearance hypothesis



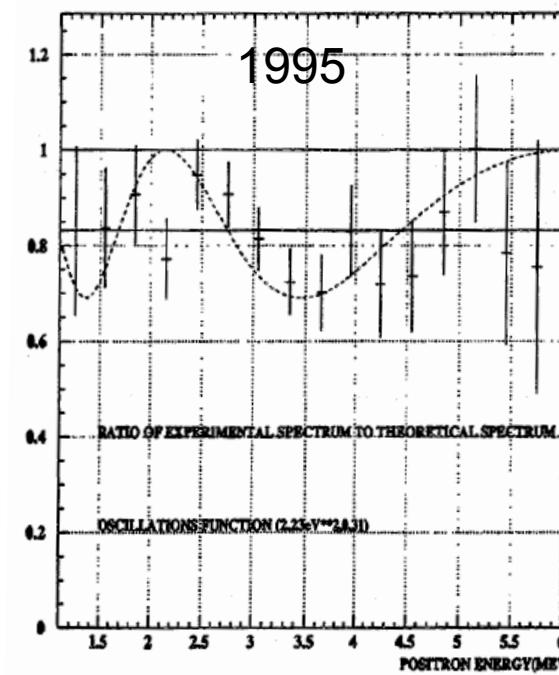
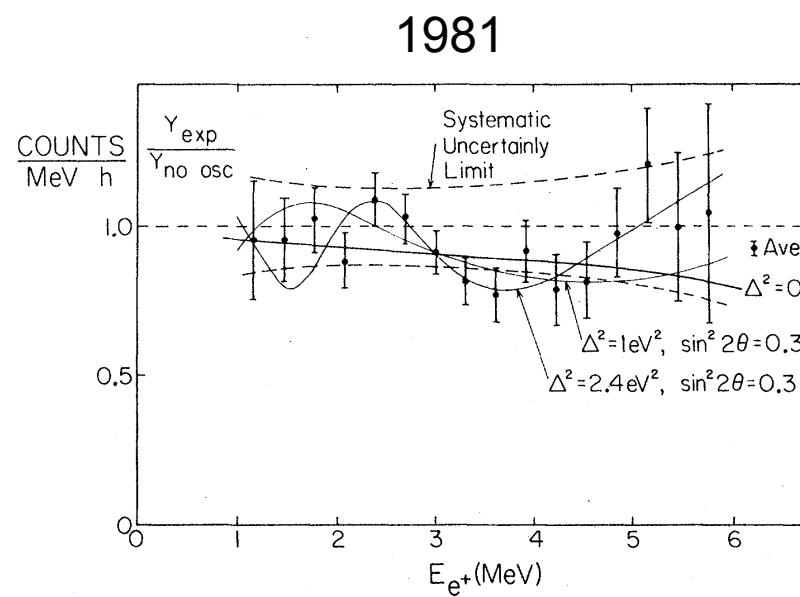
- Absence of oscillations disfavored at 98.6% C.L.

## The 1981 ILL measurement



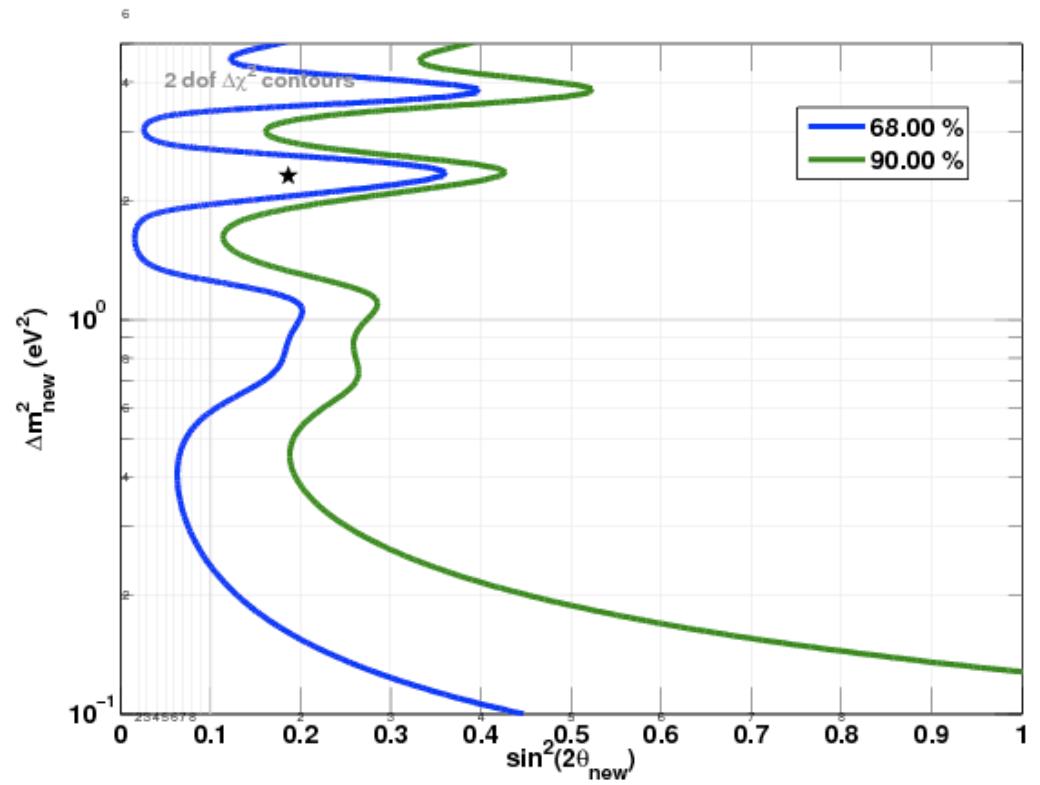
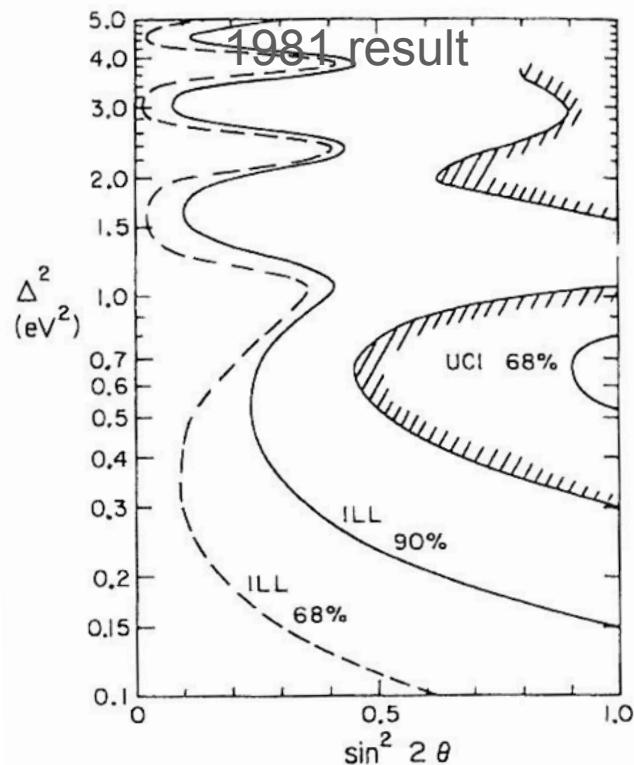
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- Reactor at ILL with almost pure  $^{235}\text{U}$ , with small core
- Detector 8 m from core
- Reanalysis in 1995 by part of the collaboration to account for overestimation of flux at ILL reactor
  - Affects the rate but not the shape analysis



Large errors, but looks like an oscillation pattern by eye ?

- 1981: Try to reproduce published contour
- 1995: Contour plot hard to follow, reproduce claim that global fit disfavors no-oscillation at  $2\sigma$
- How? Add uncorrelated systematic in each bin until it's large enough
- Quick simulation: Required error = 11%, uncorrelated, in each bin (mostly equivalent to the finite size of the reactor core in full simulation).
- We can reproduce the results quite well

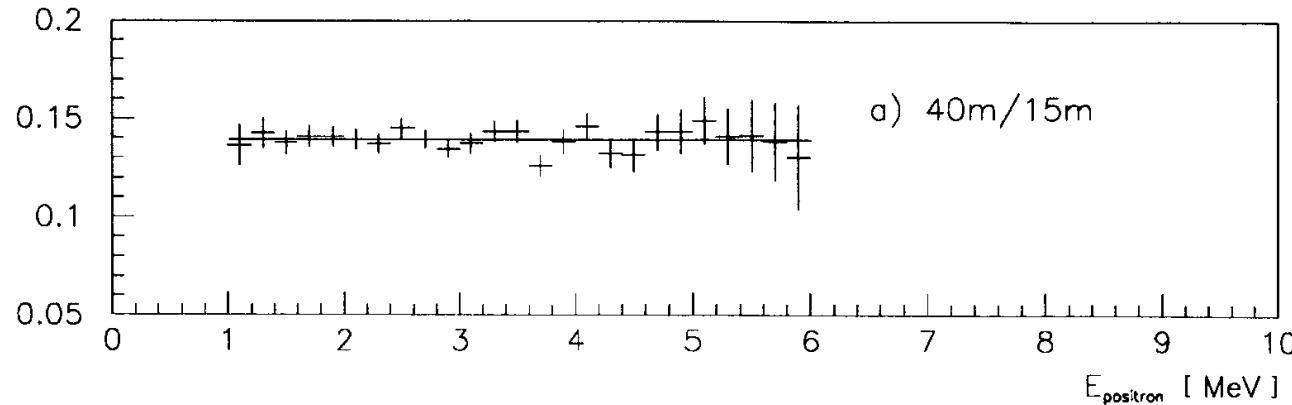


# Spectral shape analysis of Bugey-3

- Bugey-3 spectral measurements at 15 m, 40 m, 90 m
  - Best constraint from high statistics R=40 m / 15 m ratio

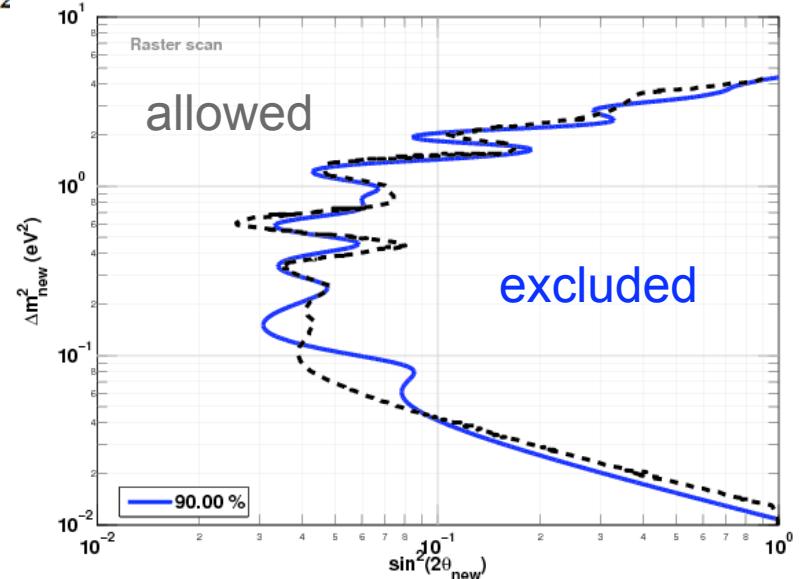
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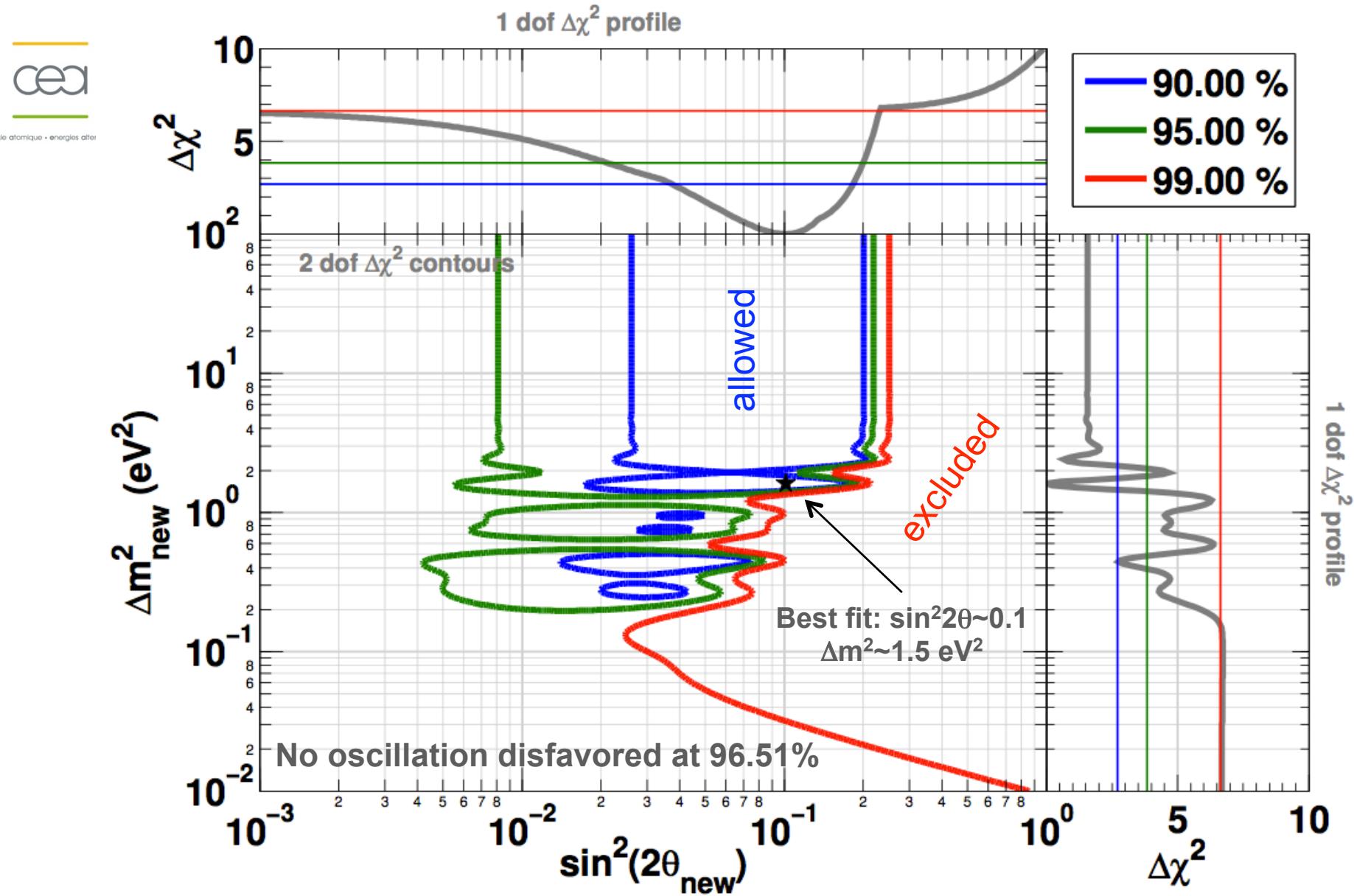


$$\chi^2 = \sum_{i=1}^{N=25} \left( \frac{(1+a)R_{th}^i - R_{obs}^i}{\sigma_i} \right)^2 + \left( \frac{a}{\sigma_a} \right)^2$$

- 2% relative systematic error
- Reproduction of the collaboration's raster-scan analysis
- Use of a global-scan in combined analysis



# Combined Reactor Rate+Shape contours

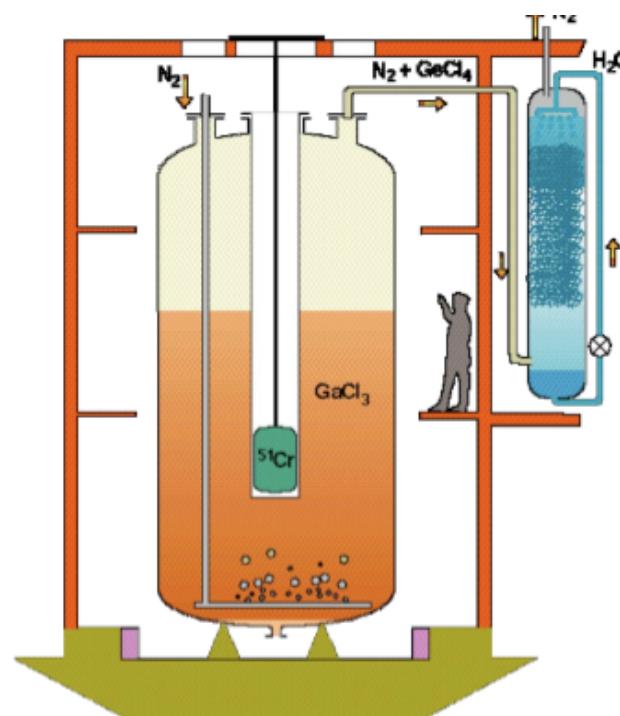
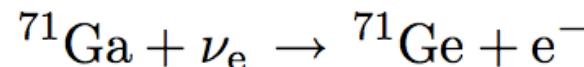


# THE GALLIUM ANOMALY

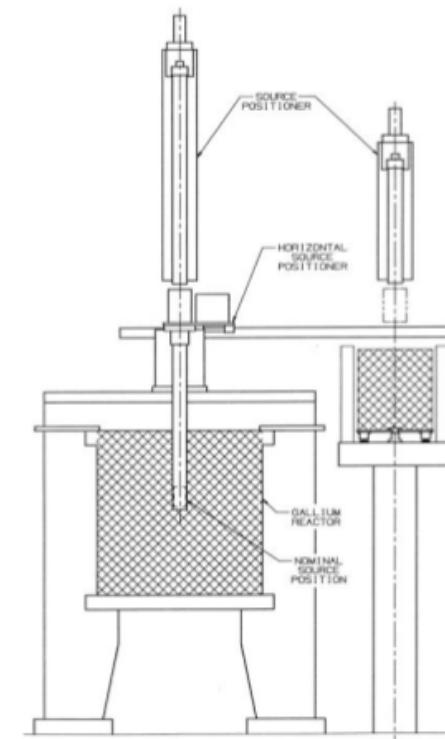
BASED ON GIUNTI & LAVEDER, PRD82 053005 (2010)

## Radiochemical experiments Gallex(left) & Sage (right)

GALLEX ( $\text{GaCl}_3$ ) and SAGE (liquid Ga) were radiochemical experiments, counting the conversion rate of  ${}^{71}\text{Ga}$  to  ${}^{71}\text{Ge}$  by (solar) neutrino capture [cannot detect anti- $\nu_e$ ]



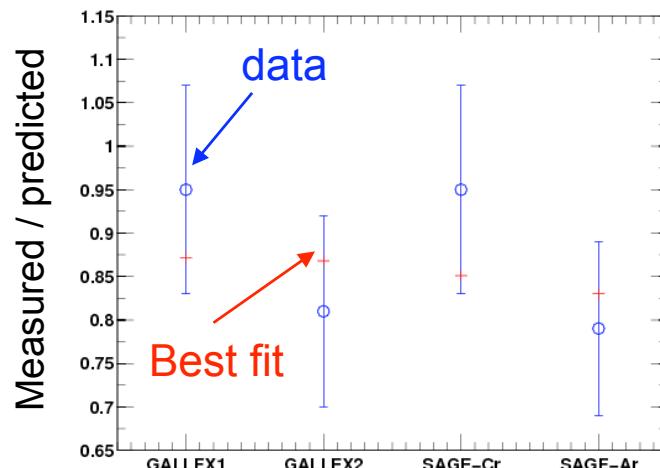
30.3 tons of Gallium  
in an aqueous solution :  $\text{GaCl}_3 + \text{HCl}$



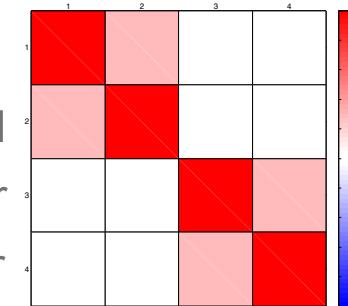
30 to 57 tons of Gallium (metal)  
In 10 tanks

# The Gallium anomaly

- 4 calibration runs with intense ( $\sim \text{MCi}$ )  $\nu_e$  (not anti- $\nu_e$ !) sources.
- Neutrinos detected through radiochemical counting of Ge nuclei:  $^{71}\text{Ga} + \nu_e \rightarrow ^{71}\text{Ge} + e^-$ 
  - 2 runs at GALLEX with a  $^{51}\text{Cr}$  source (720 keV  $\nu_e$  emitter)
  - 1 run at SAGE with a  $^{51}\text{Cr}$  source
  - 1 run at SAGE with a  $^{37}\text{Ar}$  source (810 keV  $\nu_e$  emitter)
- All observed a deficit of neutrino interactions compared to the expected activity.
- Our analysis:
  - Monte-Carlo simulation of GALLEX and SA + correlated the 2 GALLEX runs together and the 2 SAGE runs together (a bit more conservative than Giunti & Laveder PRD82 053005, 2010 to combine GALLEX & SAGE)

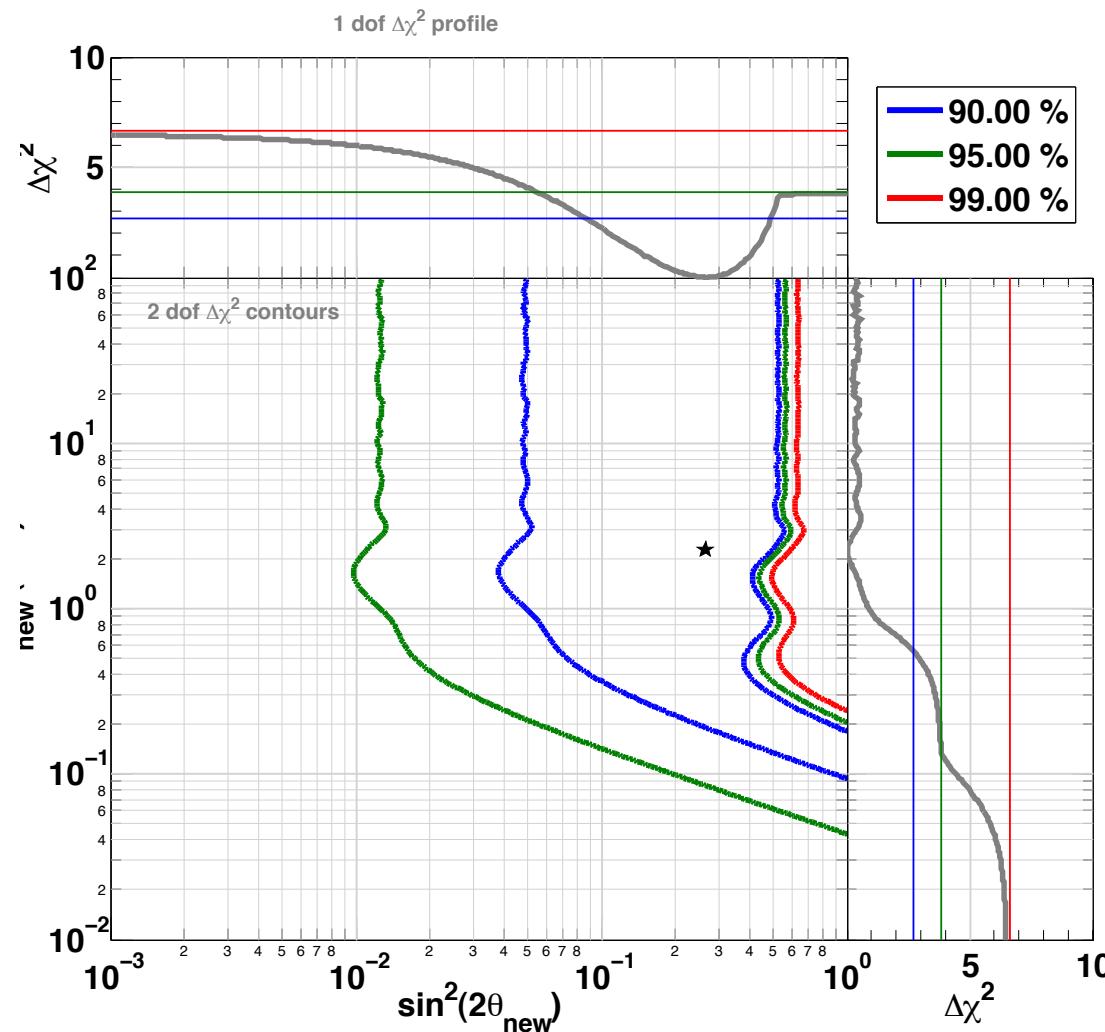


- Gallex-I
- Gallex-II
- Sage-Cr
- Sage-Ar



$$R = \text{meas./pred. rates} = 0.86 \pm 0.06 (1\sigma)$$

# The Gallium anomaly

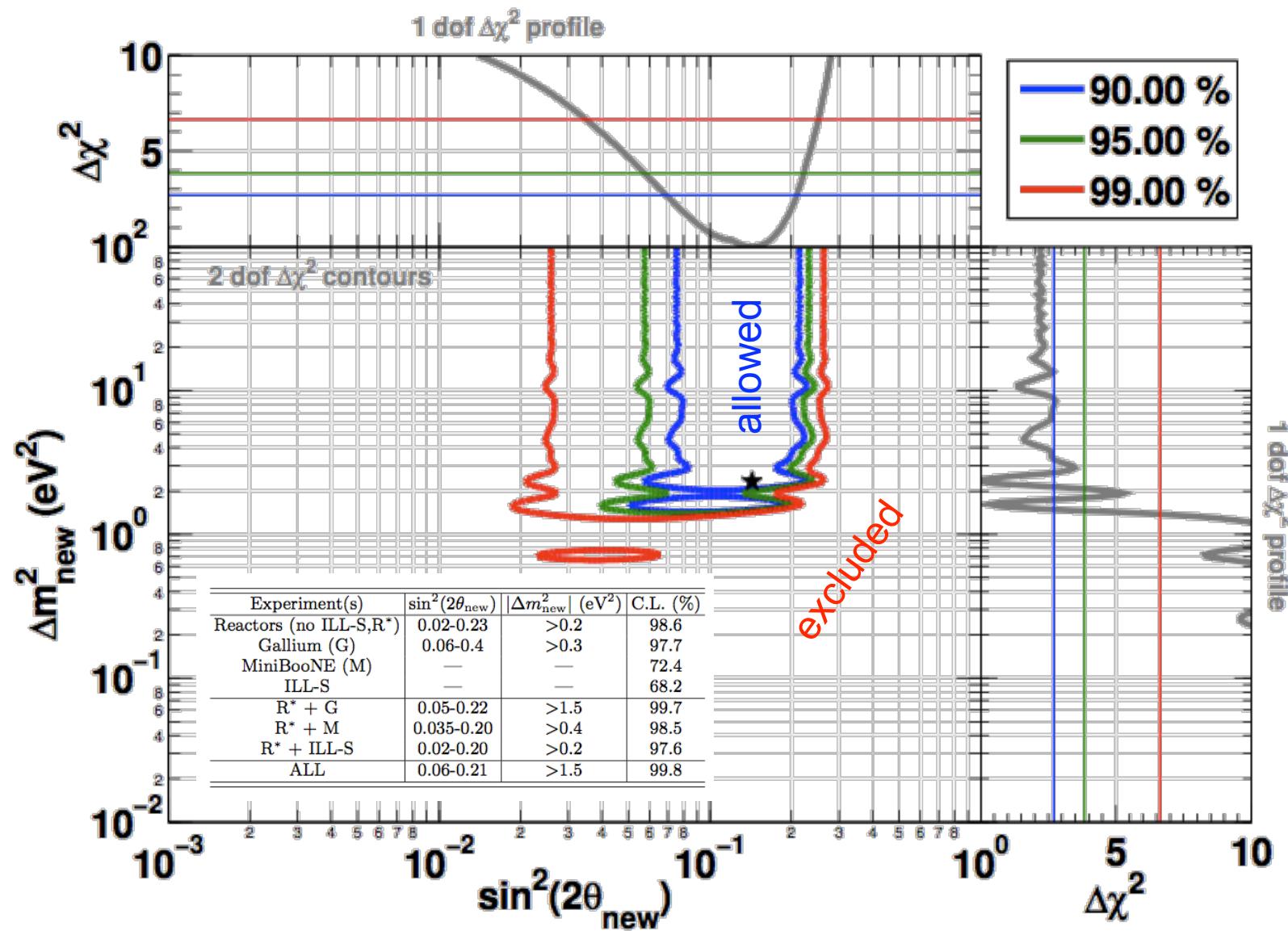


- Effect reported in C. Giunti & M. Laveder in PRD82 053005 (2010)
- Significance reduced by additional correlations in our analysis
- No-oscillation hypothesis disfavored at 97.7% C.L.

# Putting it all together: reactor rates + shape + Gallium

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The no-oscillation hypothesis is disfavored at 99.8% CL

# IMPLICATIONS FOR $\Theta_{13}$

# Implication for $\theta_{13}$ at 1-2 km baselines

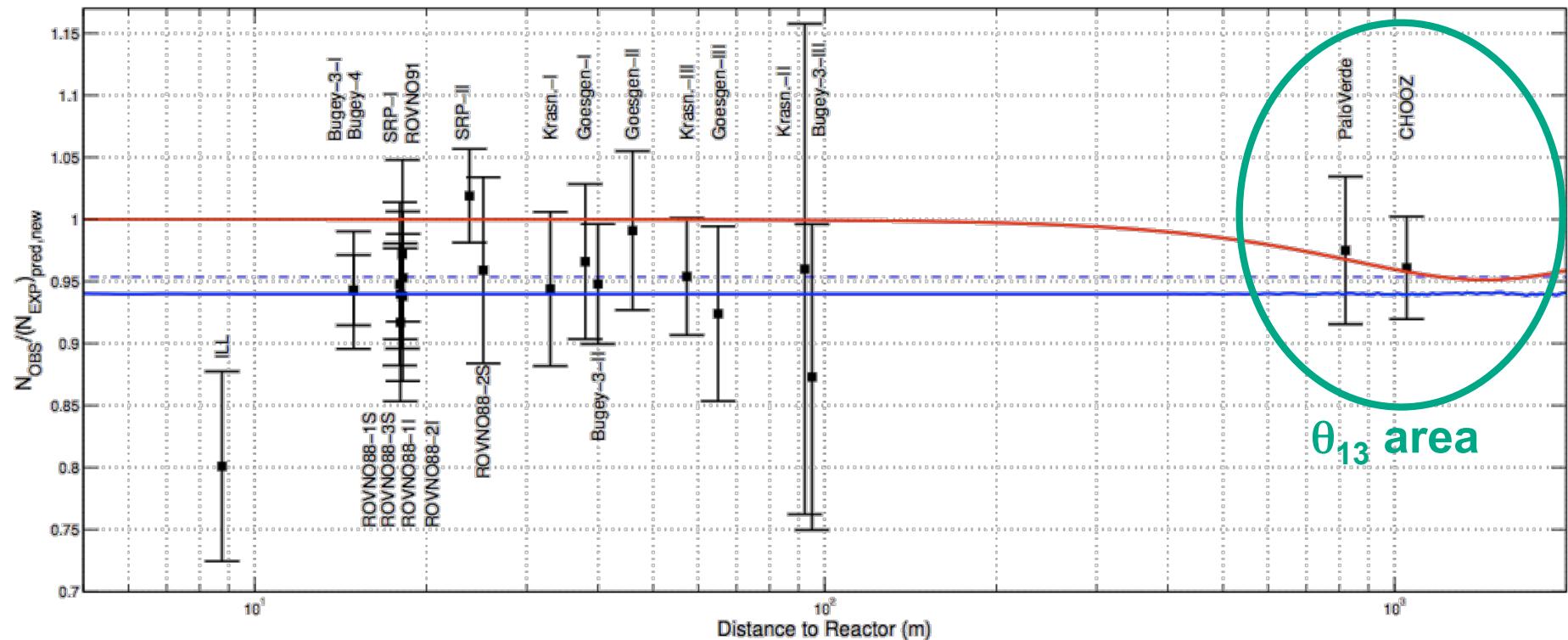


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- The choice of normalization is crucial for reactor experiments looking for  $\theta_{13}$  without near detector

$\sigma_f^{\text{pred,new}}$ : new prediction of the antineutrino fluxes

$\sigma_f^{\text{ano}}$ : experimental cross section (best fitted mean averaged)



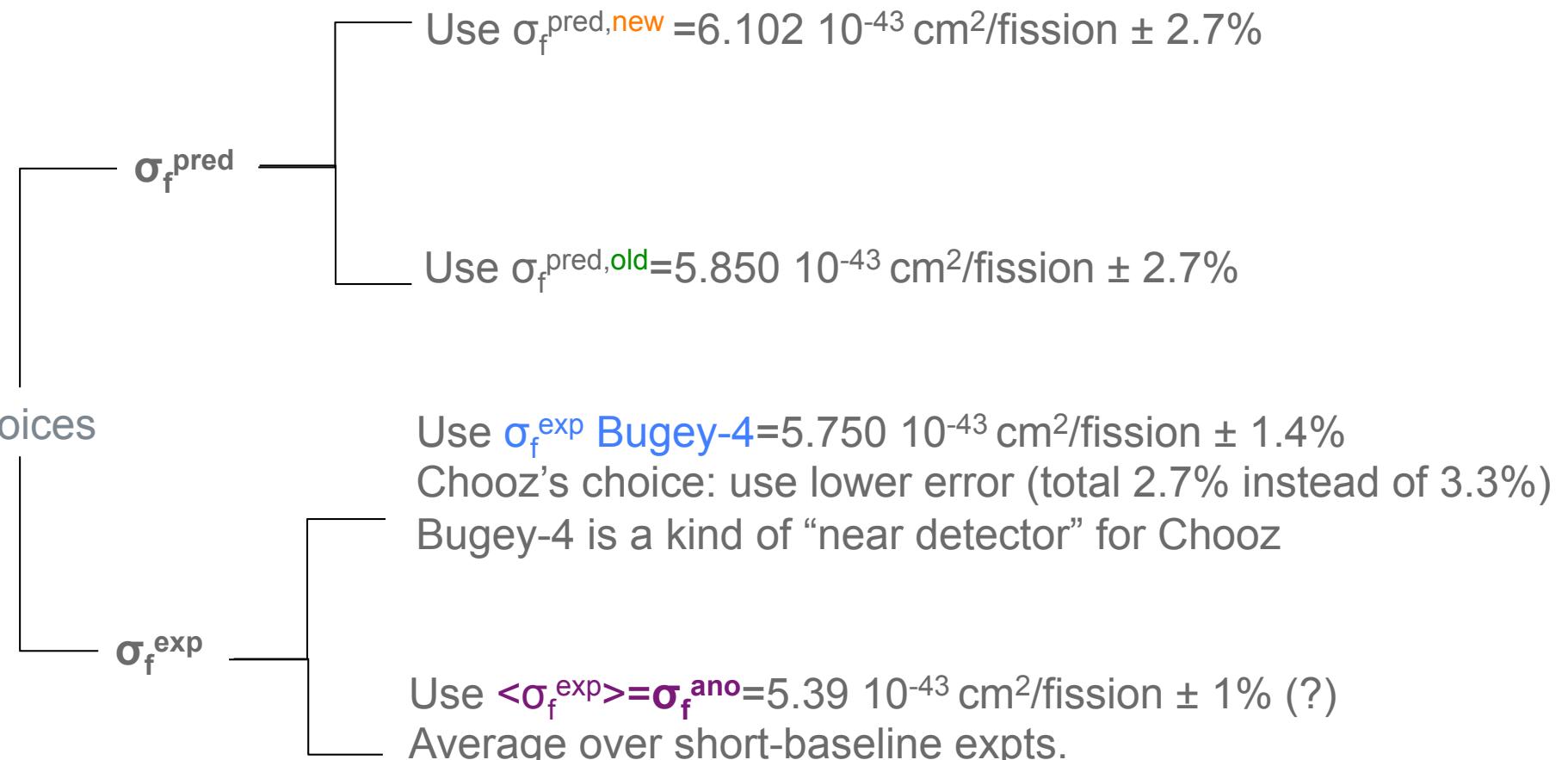
- A deficit observed at 1-2 km can either be caused by  $\theta_{13}$  or by other explanations (experimental, new physics?)

# Long baseline reactor experiments

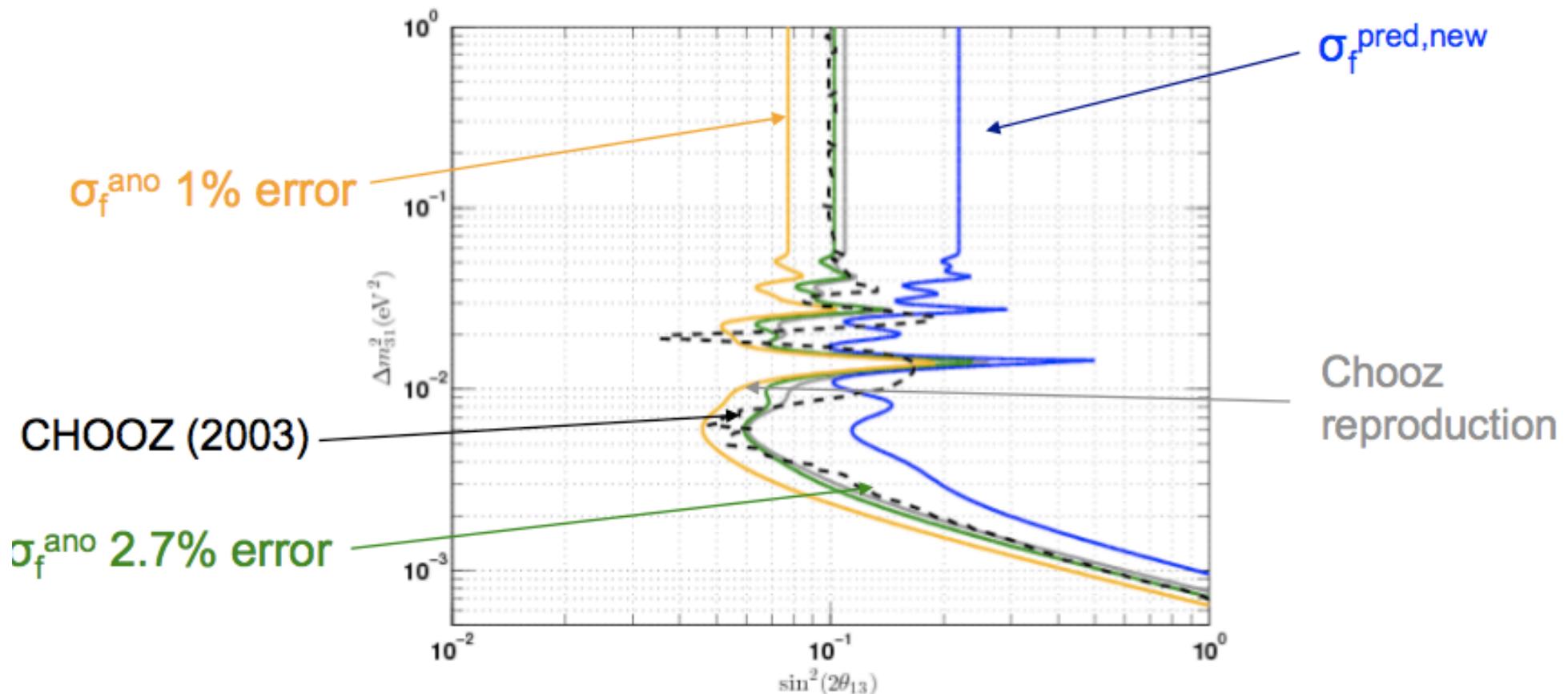


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- Experiments with baselines > 500 m
- How do you normalize the expected flux, knowing the fuel composition?  
in this slide assume Bugey-4 fuel comp.
- If **near + far detector, not an issue anymore**



- The choice of  $\sigma_f$  changes the limit on  $\theta_{13}$
- Chooz original choice was  $\sigma_f^{\text{exp}}$  from Bugey-4 with low error
- If  $\sigma_f^{\text{pred,new}}$  is used, limit is worse by factor of 2
- If  $\sigma_f^{\text{ano}}$  is used with 2.7%, we obtain the original limit  
but which error should we associate to  $\sigma_f^{\text{ano}}$ ?



# Reanalysis of KamLAND's 2010 results

arXiv:1009.4771v2 [hep-ex]

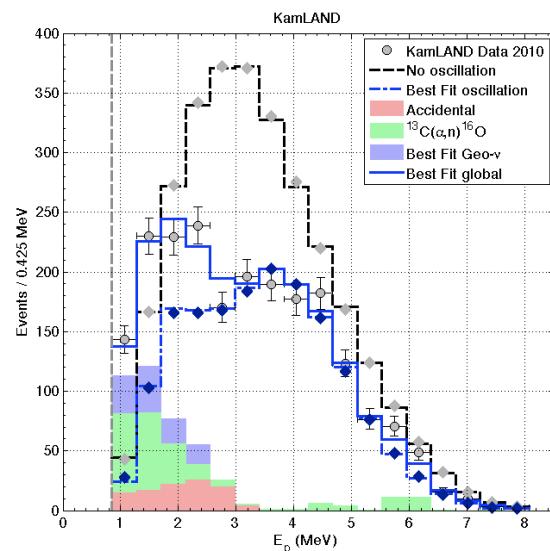


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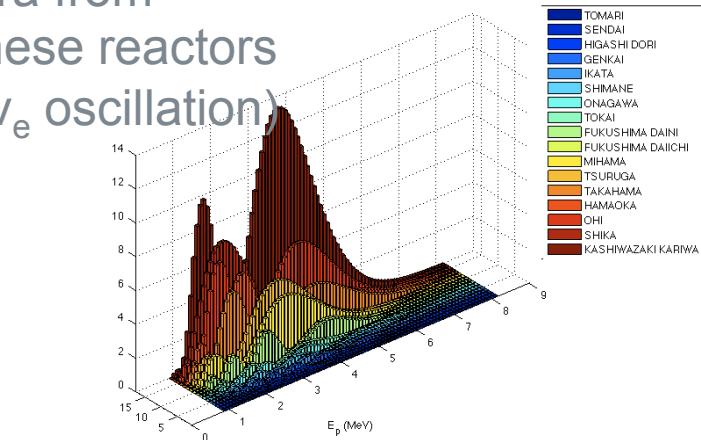
## Systematics

	Detector-related (%)	Reactor-related (%)
$\Delta m_{21}^2$	Energy scale 1.8 / 1.8	$\bar{\nu}_e$ -spectra [31] 0.6 / 0.6
Rate	Fiducial volume 1.8 / 2.5	$\bar{\nu}_e$ -spectra 2.4 / 2.4
	Energy scale 1.1 / 1.3	Reactor power 2.1 / 2.1
	$L_{cut}(E_p)$ eff. 0.7 / 0.8	Fuel composition 1.0 / 1.0
	Cross section 0.2 / 0.2	Long-lived nuclei 0.3 / 0.4
Total	2.3 / 3.0	Total 3.3 / 3.4

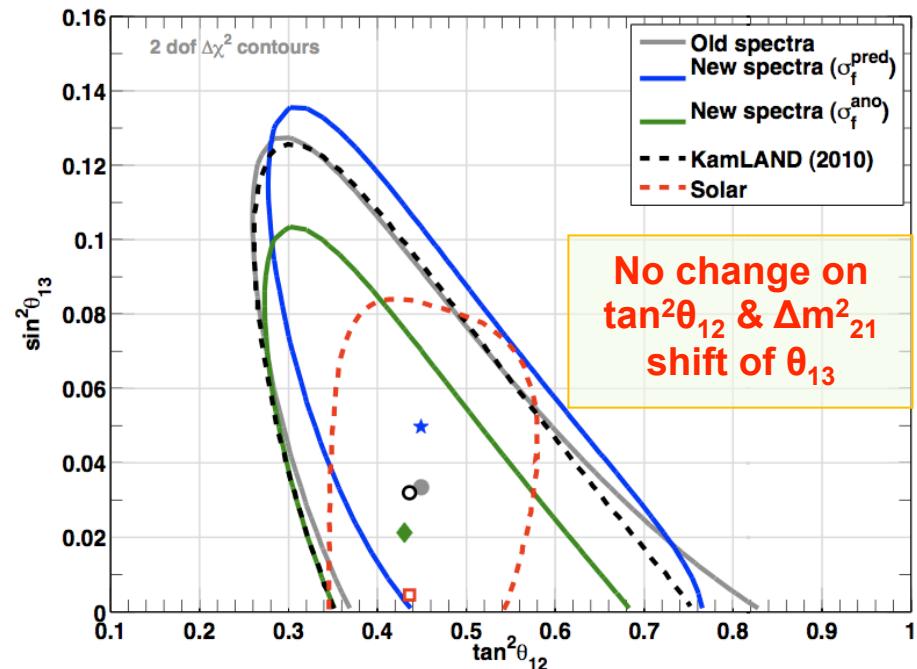
Reproduced KamLAND spectra  
within 1% in [1-6] MeV range



Spectra from  
Japanese reactors  
(with  $\nu_e$  oscillation)



With new spectra predictions

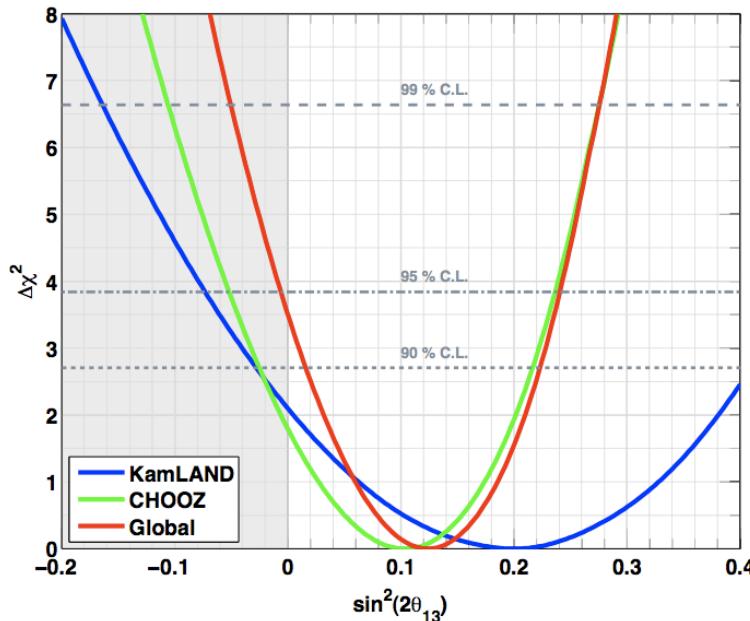


# CHOOZ and KamLAND combined limit on $\theta_{13}$



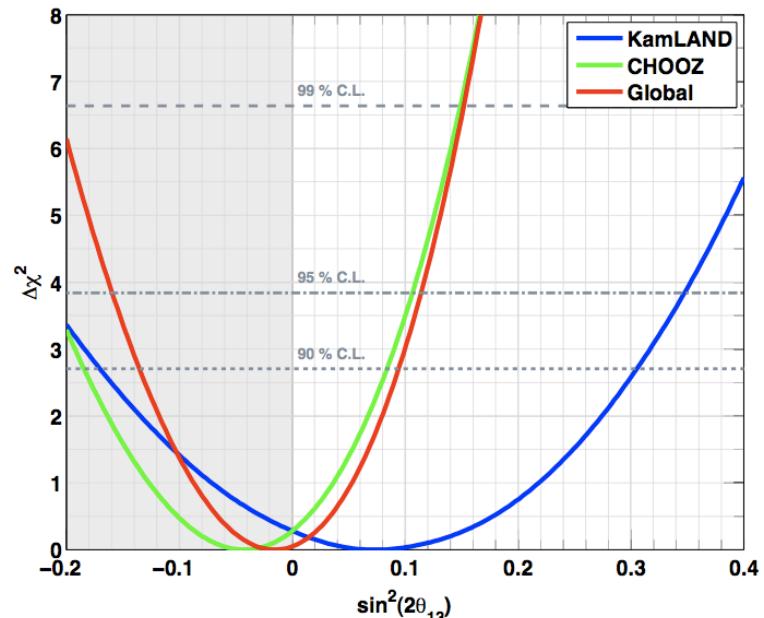
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## Normalization with $\sigma_f^{\text{pred,new}}$



use of  $\sigma_f^{\text{pred,new}}$ , 3-v framework & 2.7% uncertainty

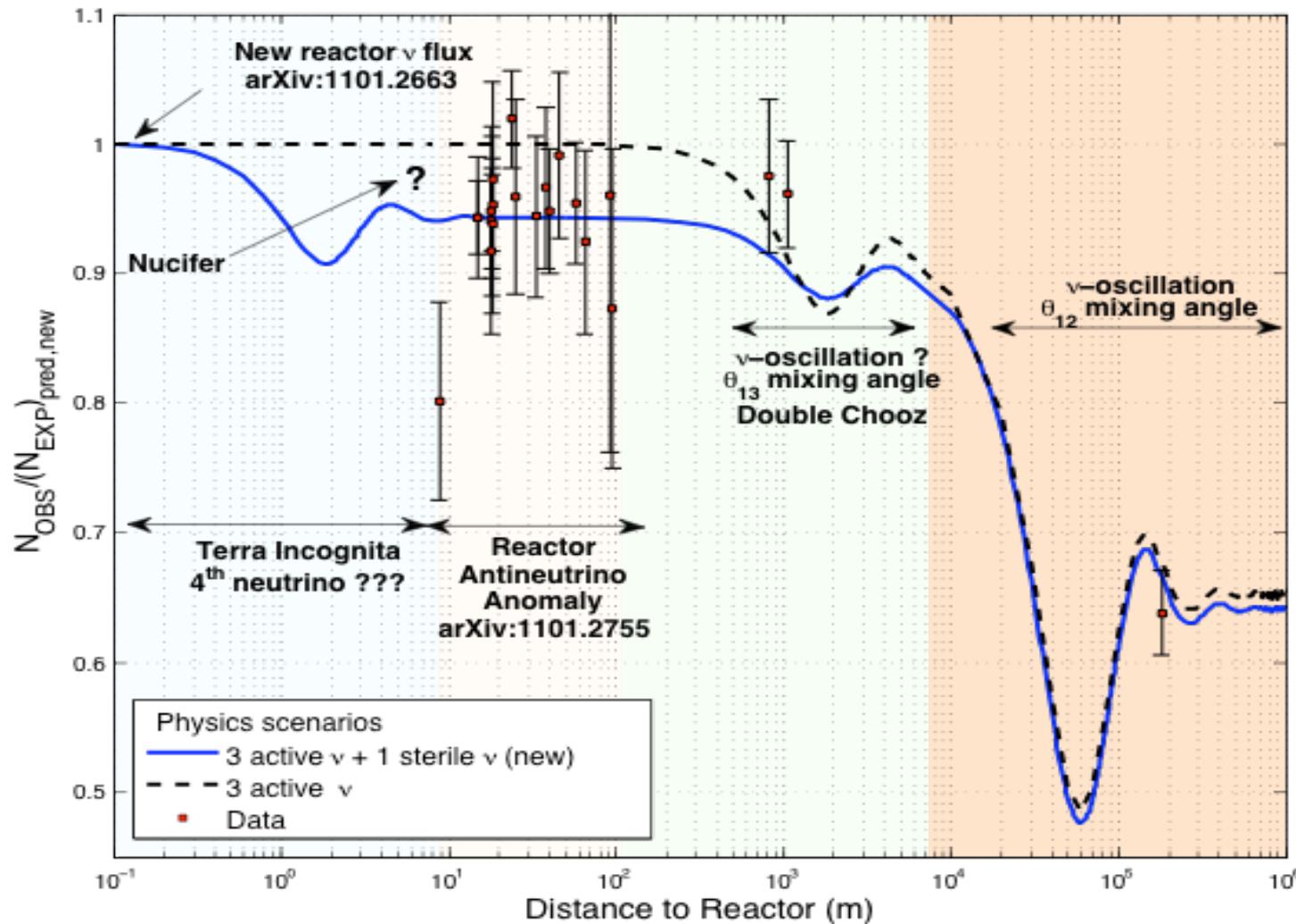
## Normalization using $\sigma_f^{\text{ano}}$



use of  $\sigma_f^{\text{ano}}$ , 3-v framework & 2.7% uncertainty (arbitrary...)

- **Our interpretation** (different from Arxiv:1103:0734 by Schwetz et al. for KamLAND)
  - No hint on  $\theta_{13} > 0$  from reactor experiments:  $\sin^2(2\theta) < 0.10$  (90% C.L., 1dof)
  - Global 90 % CL limit stays identical to published values
  - Multi-detector experiments are not affected

## The 4th Neutrino Hypothesis: need new experimental inputs !



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## New Reactor Antineutrino Anomaly Discovered

- Experimental bias to be deeply investigated
- New physics hypothesis tested: 4<sup>th</sup> neutrino
  - no-oscillation hypothesis disfavored at 99.8%

**Clear experimental confirmation / infirmation is needed:**

- $L/E \approx$  few m/MeV or km/GeV

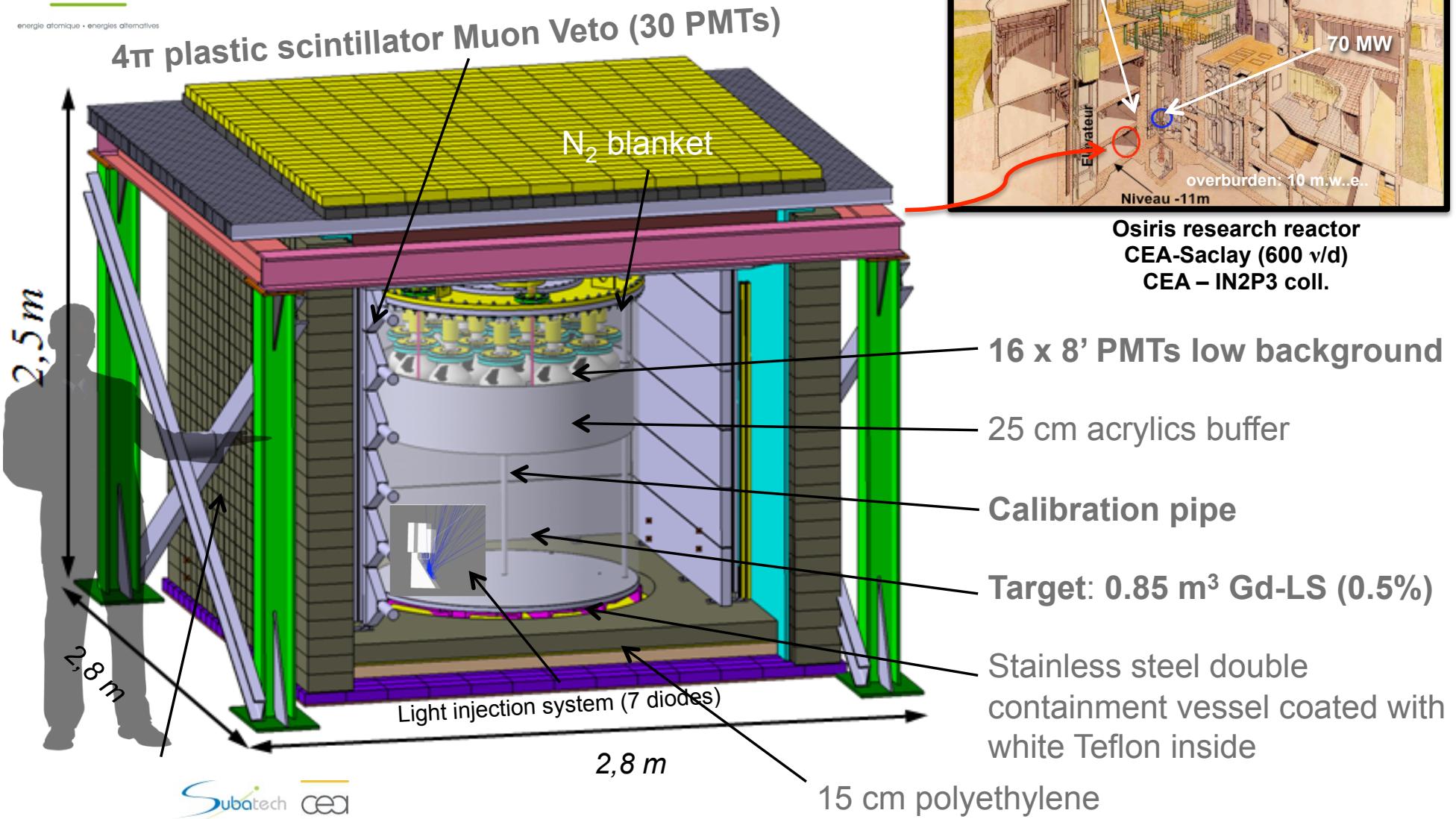
### New Experiment at Reactor

Short Baseline – Shape + Rate Analysis: SCRAAM, Nucifer,...

**Mci neutrino generator in/close to a large liquid scintillator**  
Like SNO+, Borexino, KamLAND

**New neutrino beam experiment probing for electron GeV neutrino disappearance at 100 m & 1 km**  
C. Rubbia's proposal at CERN-PS

...

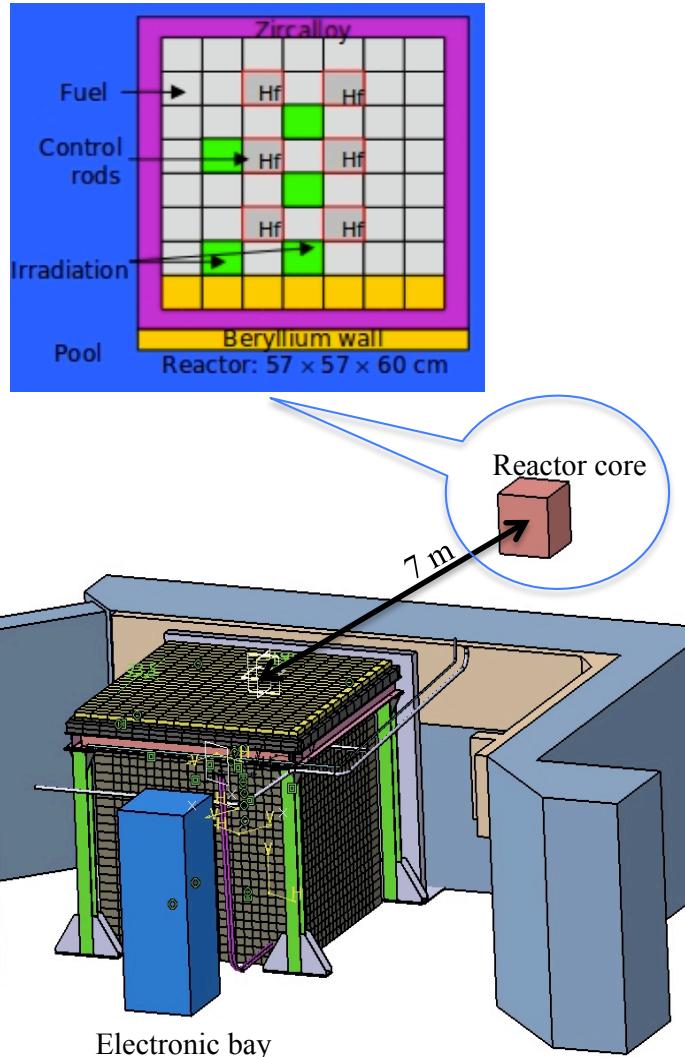




# The nuclear core compactness

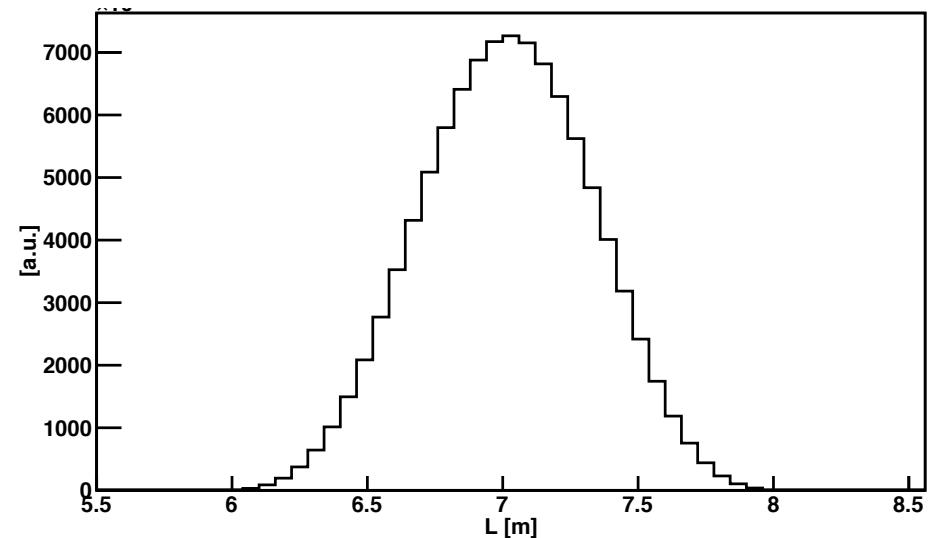
cea

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- Core Size:  $57 \times 57 \times 60 \text{ cm}$
- Detector Size :  $1.2 \times 0.7 \text{ m}$  (850l)
- baseline distribution
  - $\langle L \rangle = 7.0 \text{ m}$
  - variance :  $0.3 \text{ m}$
  - eV<sup>2</sup> oscillations are not washed out

Baseline distribution

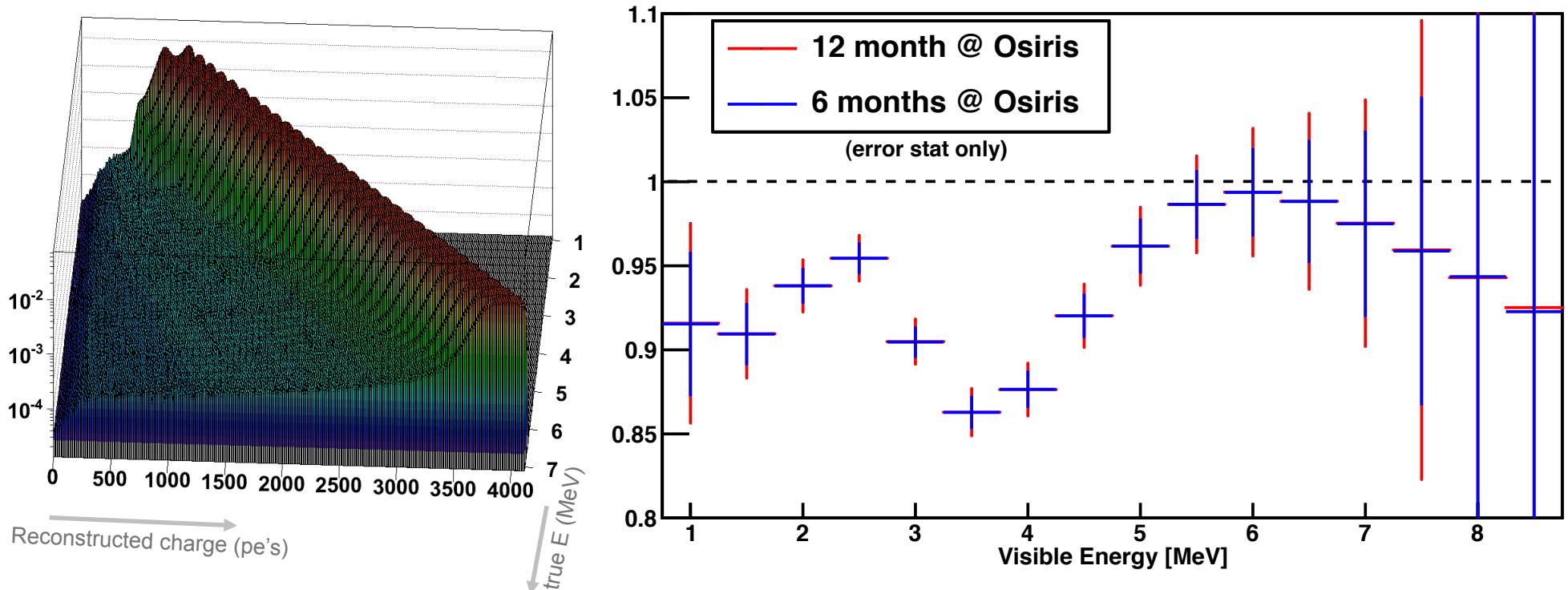


# Nucifer attempt testing the anomaly



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- Folding the Nucifer Geant4 Monte Carlo detector response
  - Energy resolution from Geant4 simulation (not fully tuned yet)
  - Statistical error bars for 12 & 24 months of data at Osiris
  - $\Delta m^2 = 2.4 \text{ eV}^2$  &  $\sin^2(2\theta)=0.15$
  - No backgrounds. Thus to be taken with a grain of salt ...

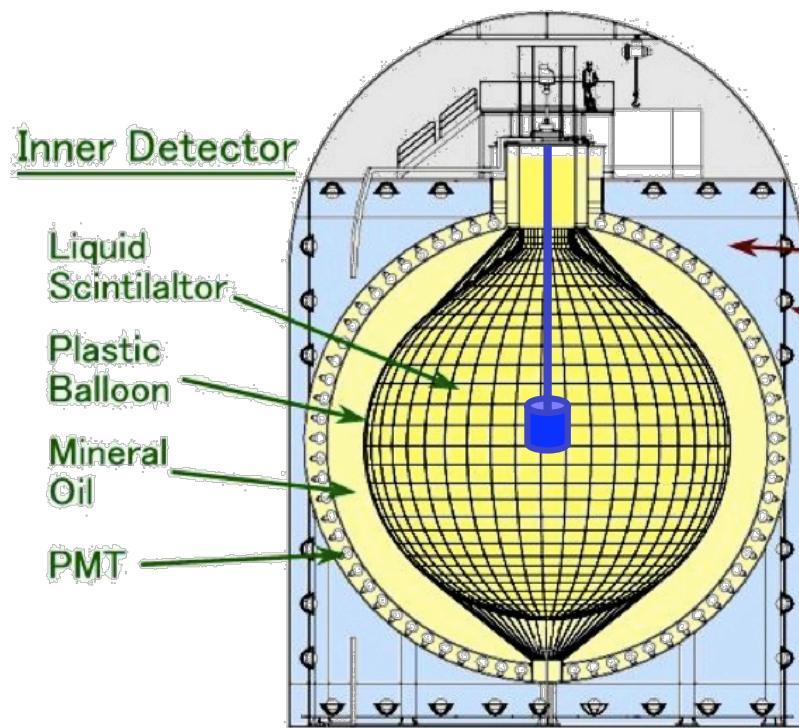


# MCi $^{51}\text{Cr}/^{37}\text{Ar}$ Experiment Concept

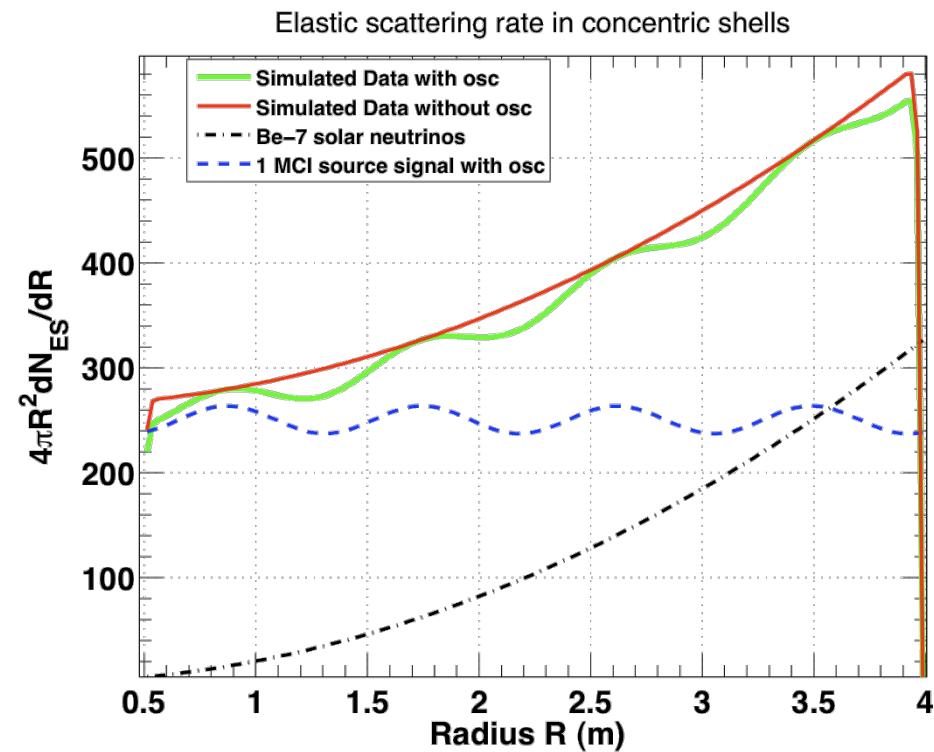


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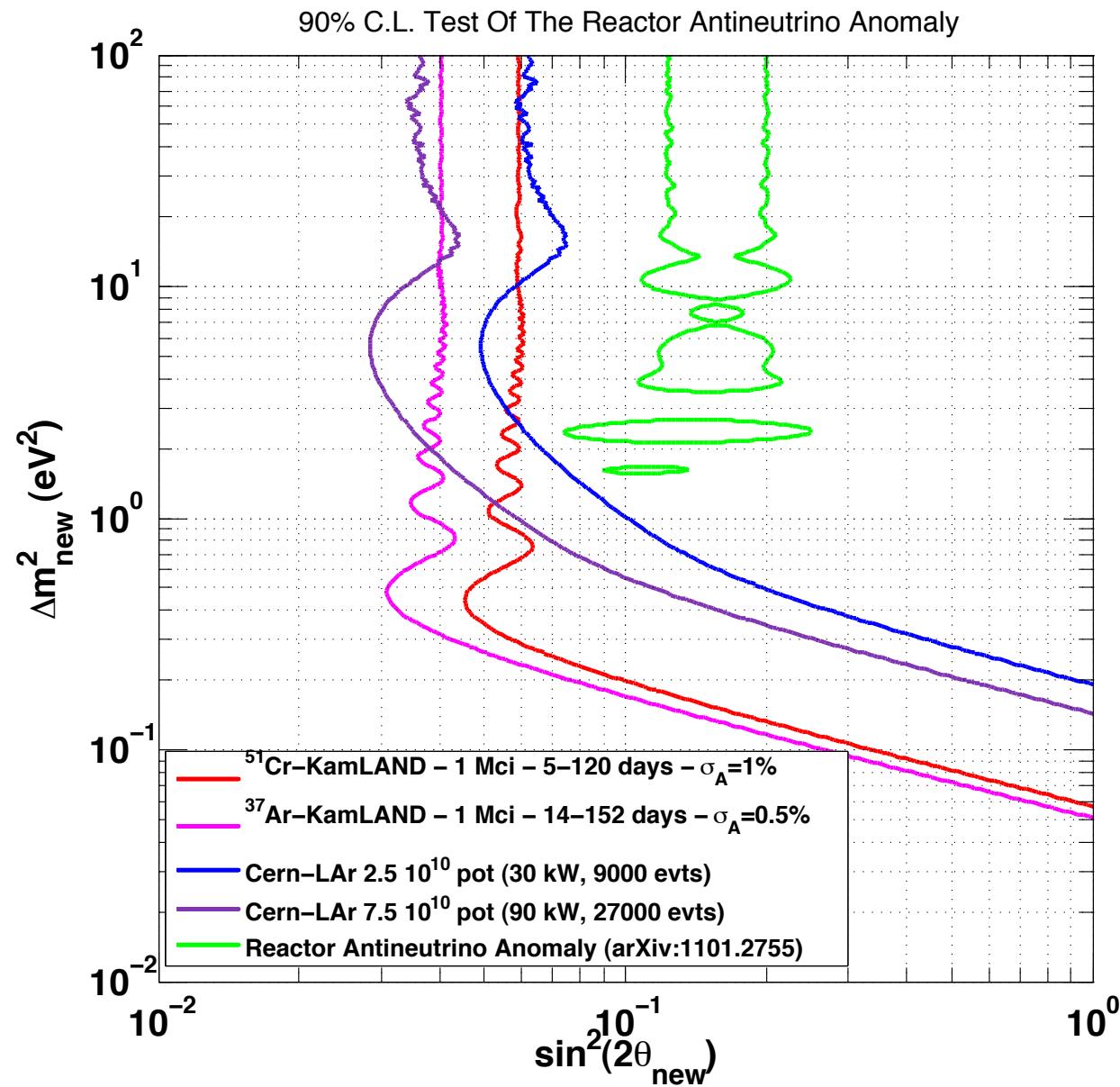
- A strong 1 MCi  $\nu_e$  source in the middle of a large LS detector
- Elastic scattering on  $e^-$  (few 1,000 evts, 150 days, >250 keV)
- A good resolution in position (20 cm)



Real oscillation pattern VS radius

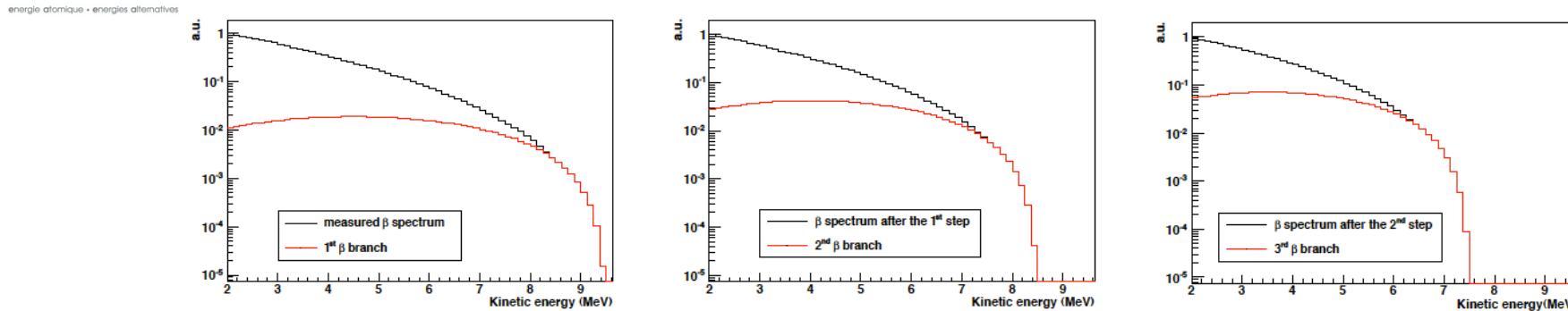


# Promising experimental prospect testing the RAA!

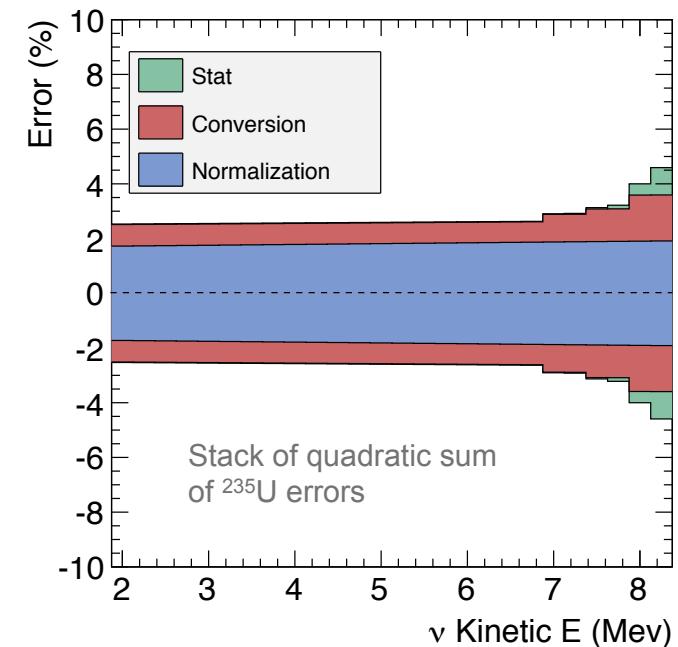


# BACKUP SLIDES

Lost info of single  $\beta$ -branches  $\rightarrow$  fit  $e^-$  (50 keV bins) spectrum with a sum of 30 effective branches



- All theory included in these effective branches but:
  - What  $Z$ ? : Mean fit on nuclear data  $Z=f(E_0)$
  - $Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, \quad Z \geq 34$
  - What  $A_{CW}$ ? : effective correction
  - $\Delta N_\nu^{C,W}(E_\nu) \approx 0.65 \times (E_\nu - 4 \text{ MeV}) \quad \%$
- Conversion error from envelop of all numerical studies

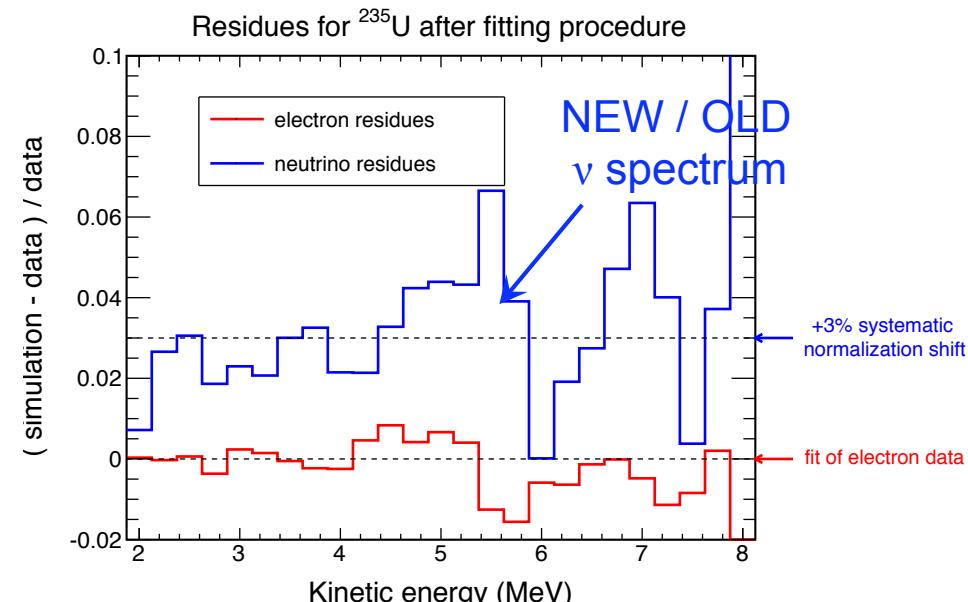
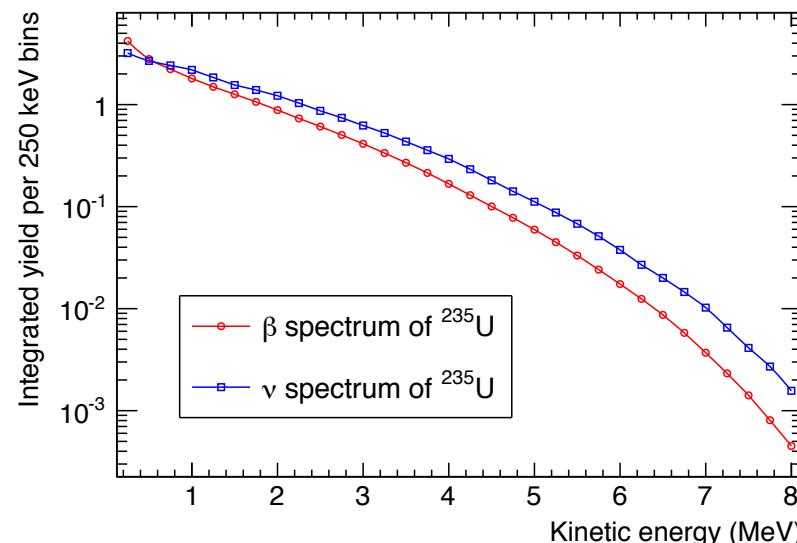
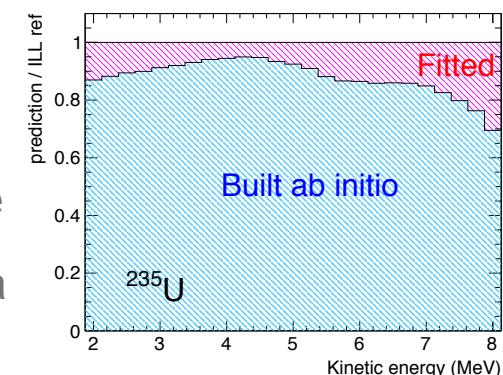


# New Saclay-Subatech reactor $\nu$ -spectra from ILL $\beta$ -spectra

[arXiv:1101.2663 \[hep-ex\]](https://arxiv.org/abs/1101.2663), submitted to PRC



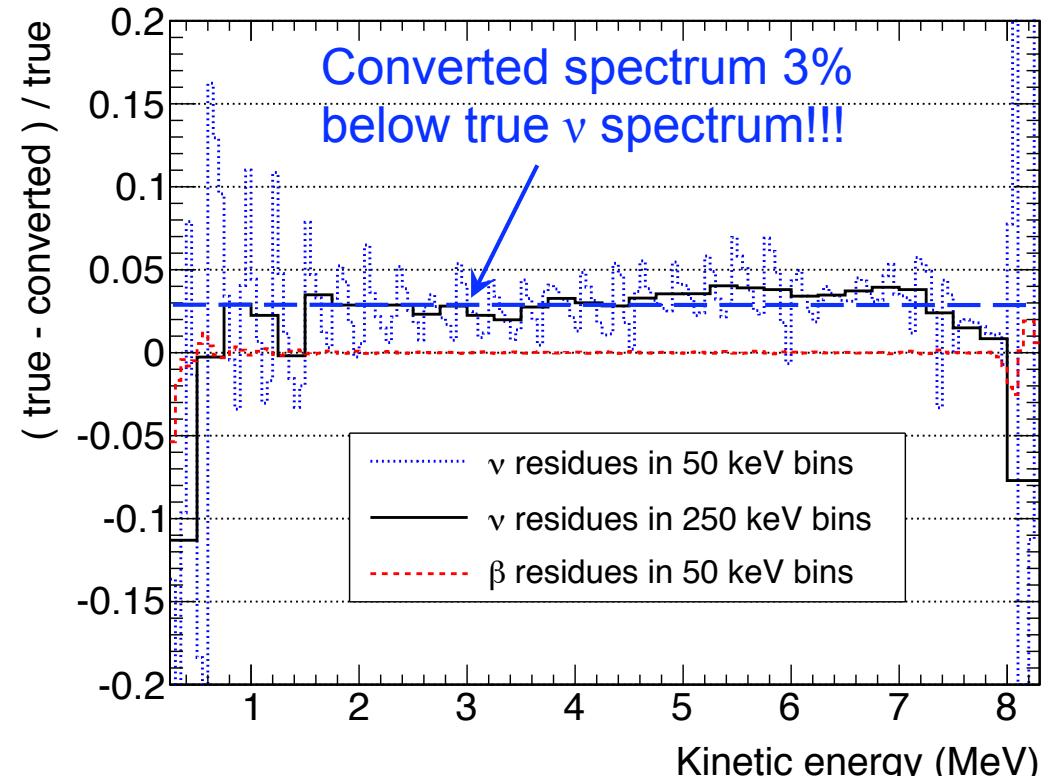
- Starting point: same  $\beta$ -spectra from measurement at the ILL reactor in the 80's
- Conversion from electron to anti- $\nu_e$  spectra:
  - OLD: **30 effective branches** method
  - NEW: conversion method taking into account the **whole information of nuclei** measured and stored in **nuclear databases** (~90% info from data bases, ~10% fitted with **5 effective branches**)
- Full Error Propagation and Correlations included



- Result: **+3% bias** (averaged) with respect to previous results
- Similar results for all measured isotopes ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ )

## Stringent test

1. Define “true”  $e^-$  and  $\nu$  spectra from reduced set of well-known branches from ENSDF nuclei data base.
2. Apply exact same **OLD** conversion procedure to true  $e^-$  spectrum.
3. Compare the converted  $\nu$  spectrum to the true one.
4. This technique gives a 3% bias compared to the true  $\nu$  spectrum



=> The **OLD** effective conversion method biases the predicted  $\nu$  spectrum at the level of -3% in normalization.

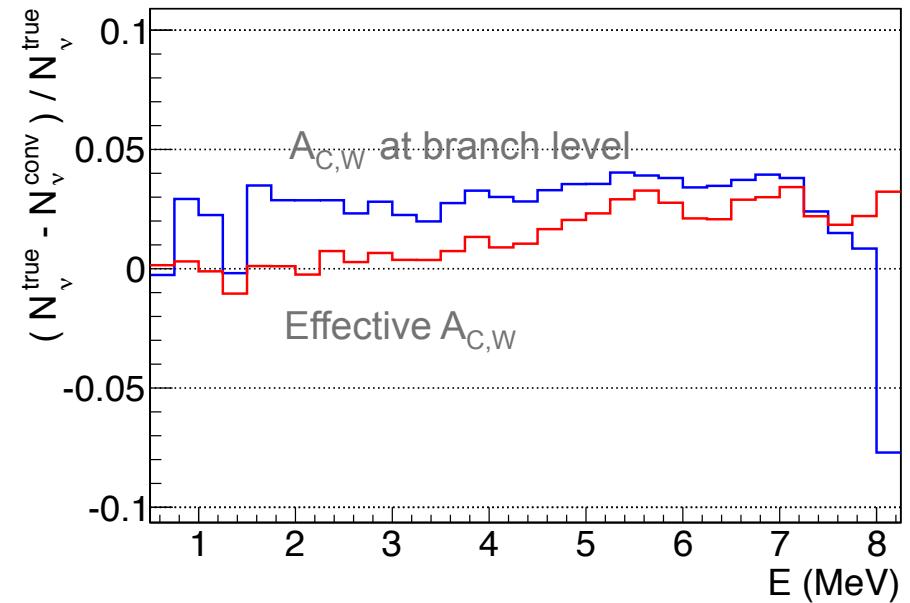
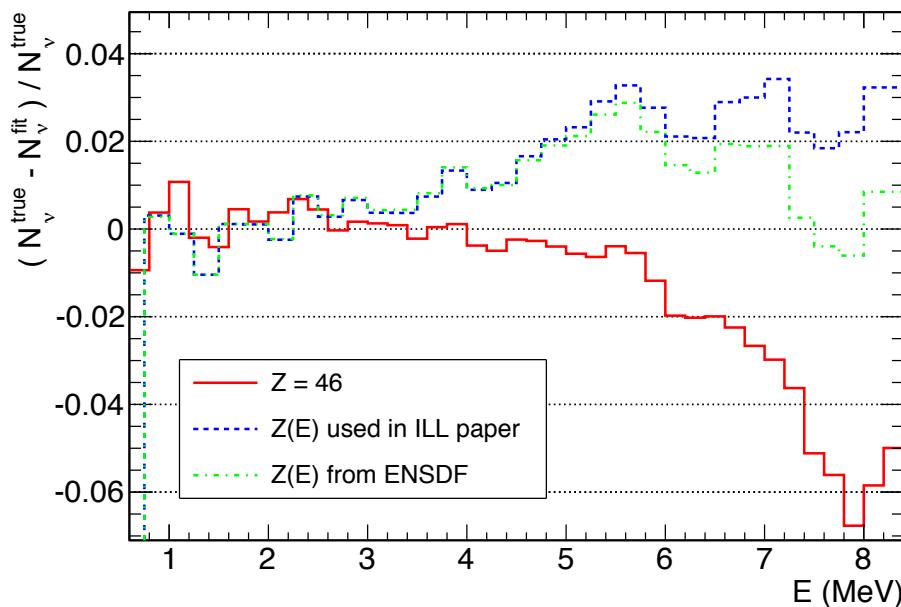
# Origin of the 3% shift



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- $E < 4$  MeV: deviation from effective linear  $A_{C,W}$  correction of ILL data

$$\Delta N_\nu^{C,W}(E_\nu) \approx 0.65 \times (E_\nu - 4 \text{ MeV}) \quad \%$$



- $E > 4$  MeV: mean fit of  $Z(E_0)$  doesn't take into account the very large dispersion of  $Z$  around the mean curve

$$Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, \quad Z \geq 34$$